

# 21 cm Power Spectrum Experiments

UC Irvine, Oct. 1, 2015

Aaron Parsons

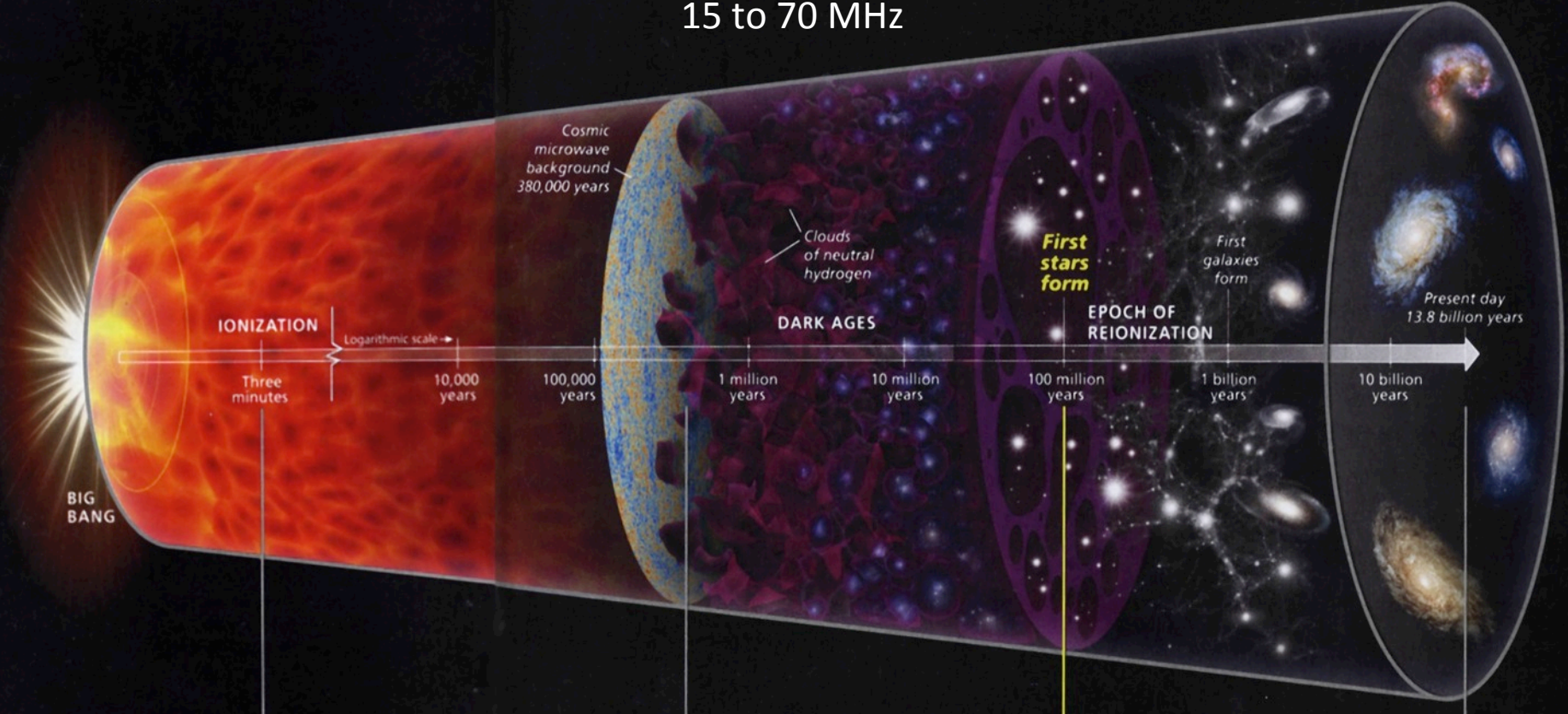


# 21-Centimeter Cosmology Explained

$z = 100$  to  $20$   
15 to 70 MHz

$z = 20$  to  $5$   
70 to 240 MHz

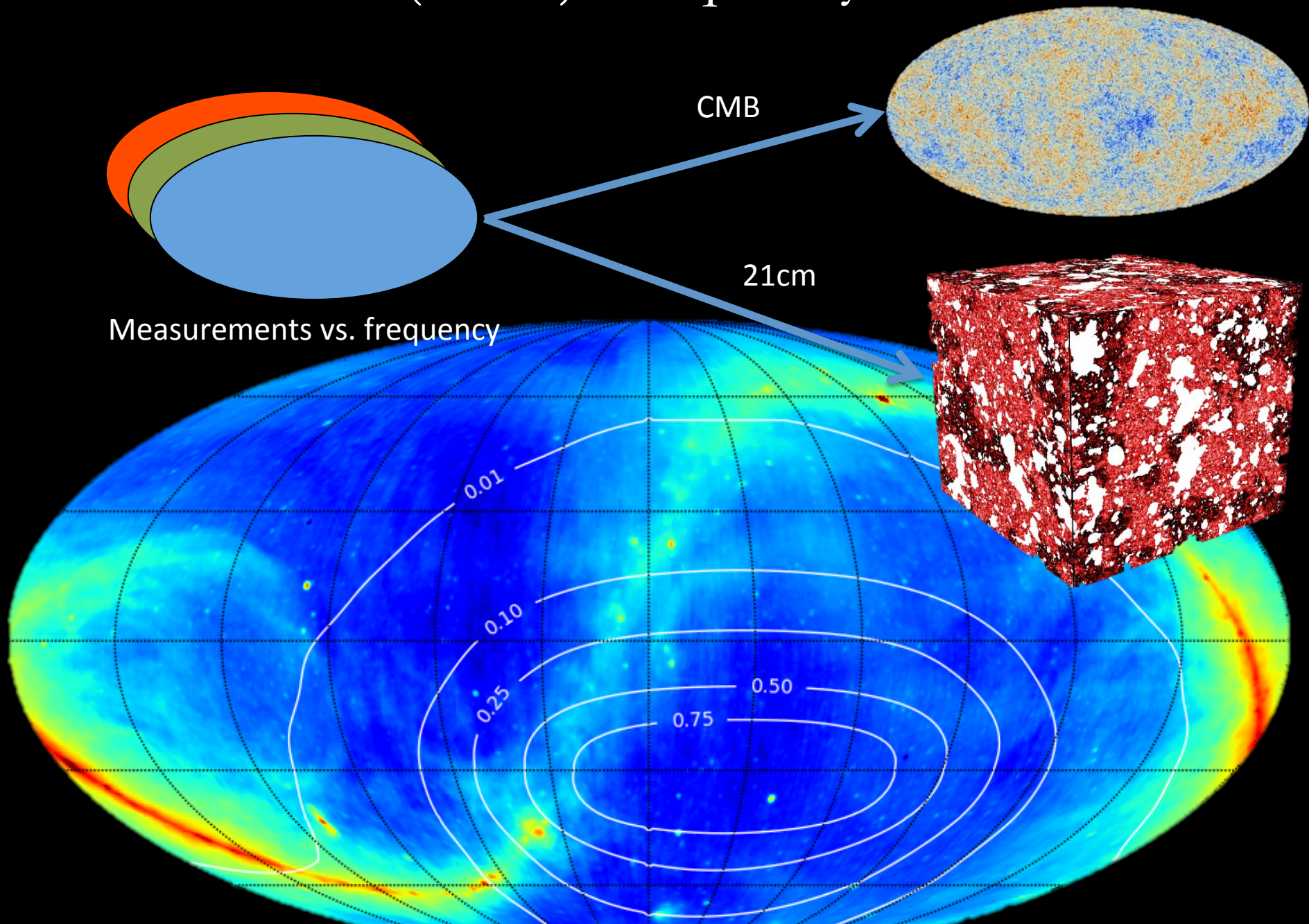
$z = 5$  to  $0$   
240 to 1400 MHz

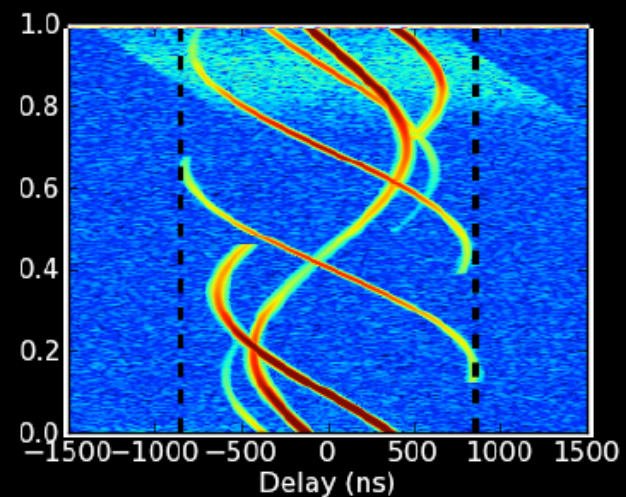
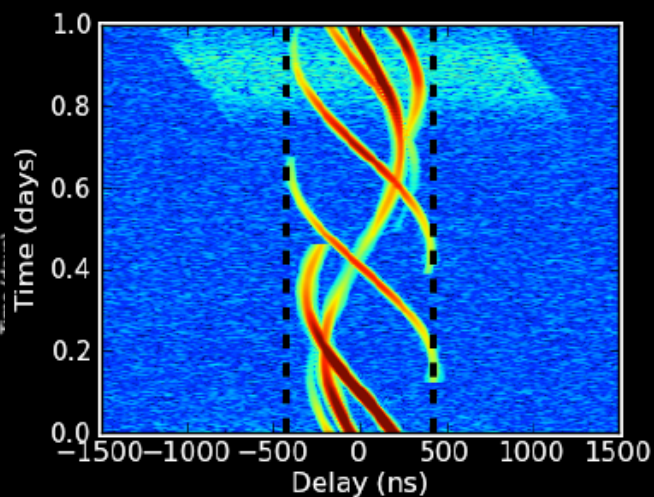
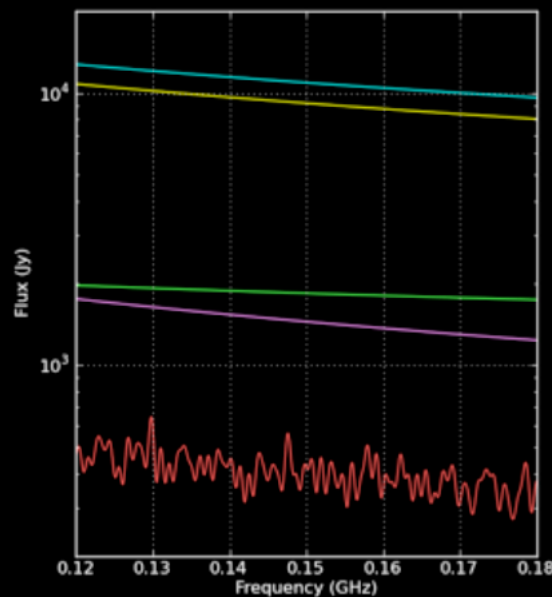
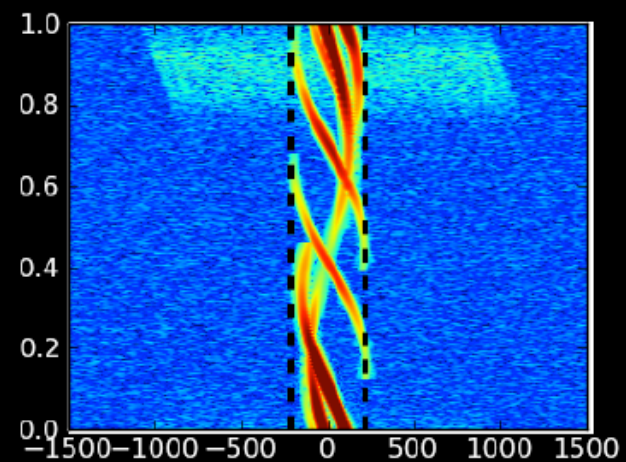
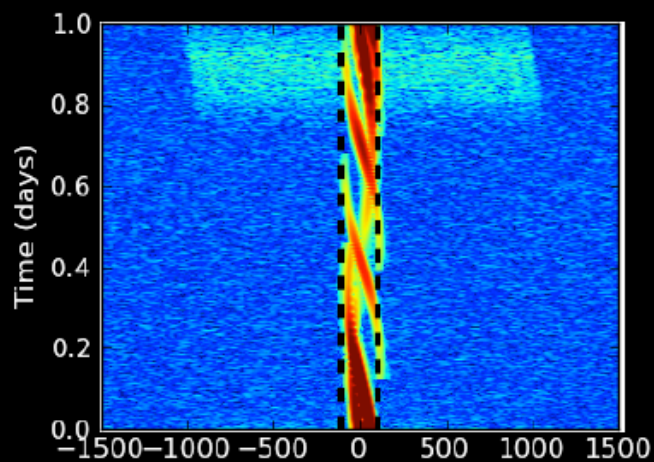
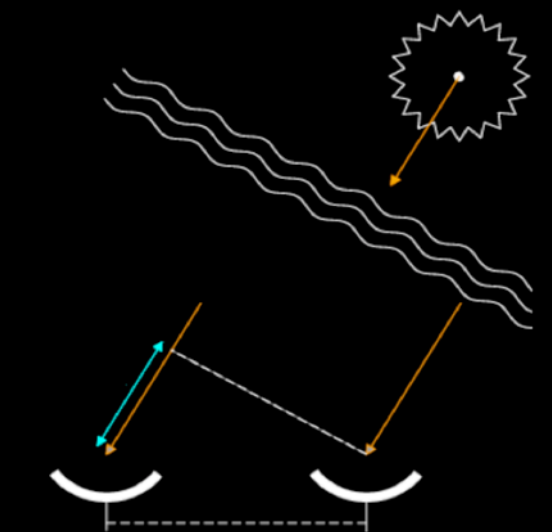


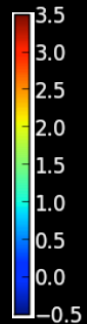
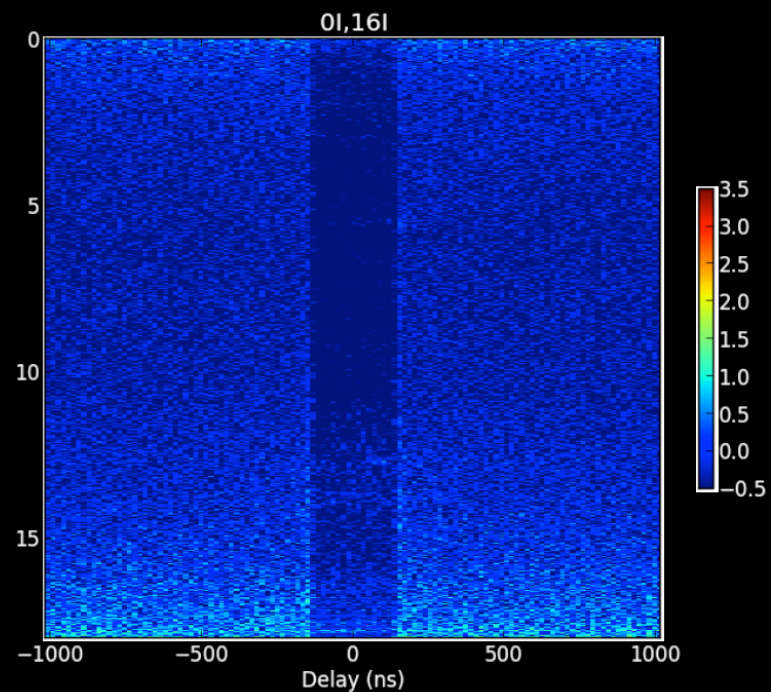
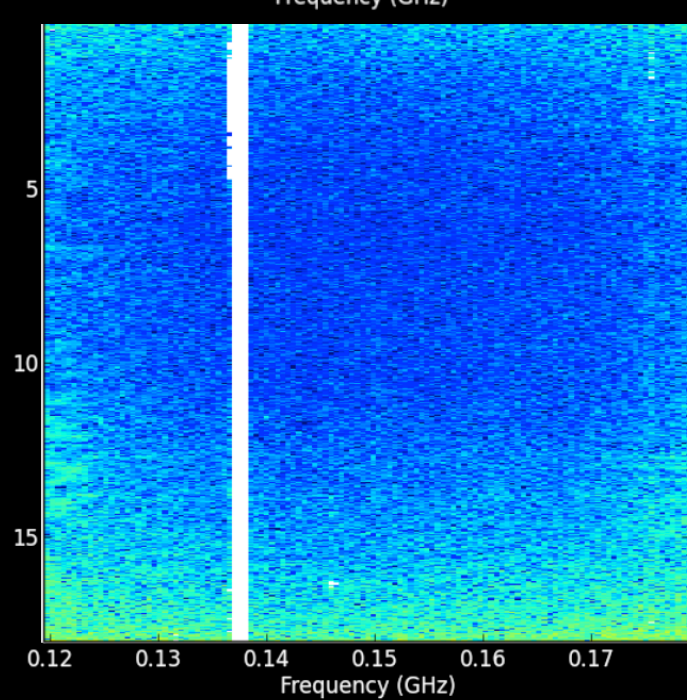
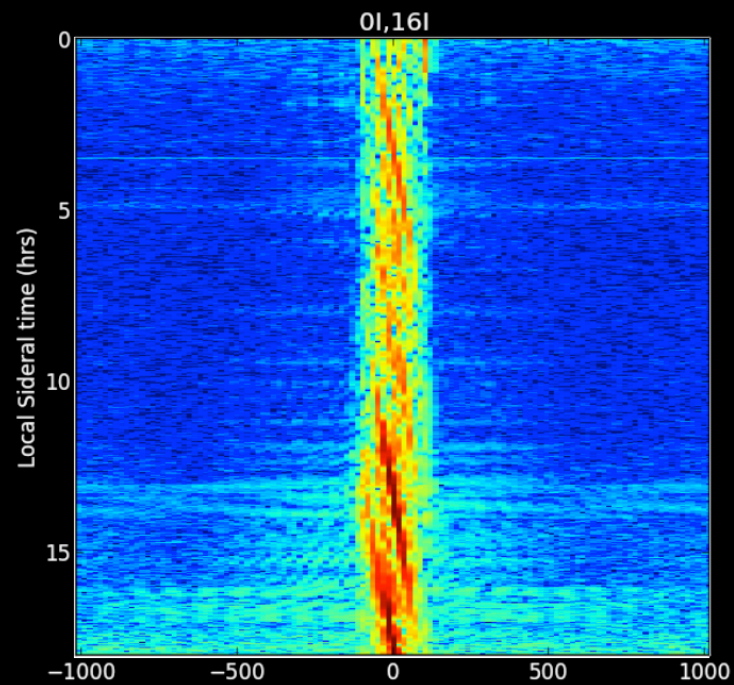
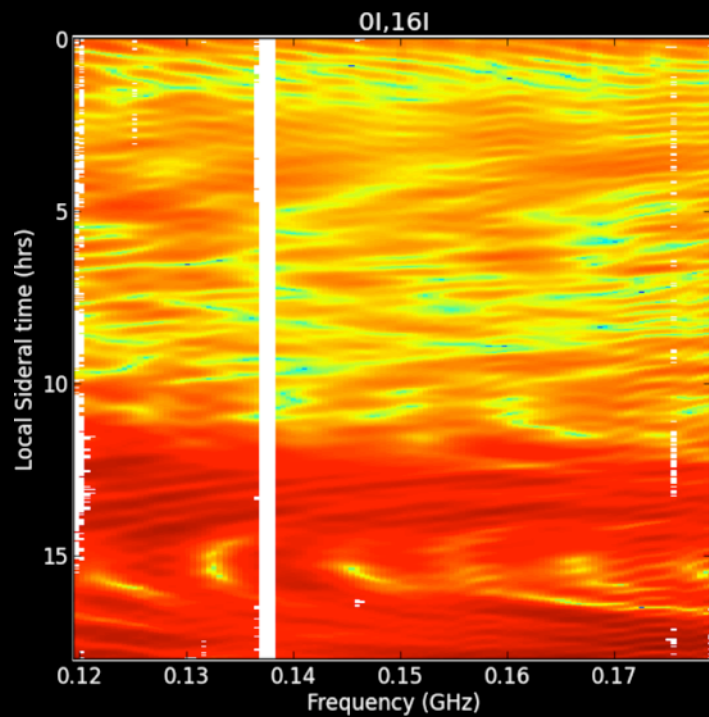
# Why 21cm Cosmology

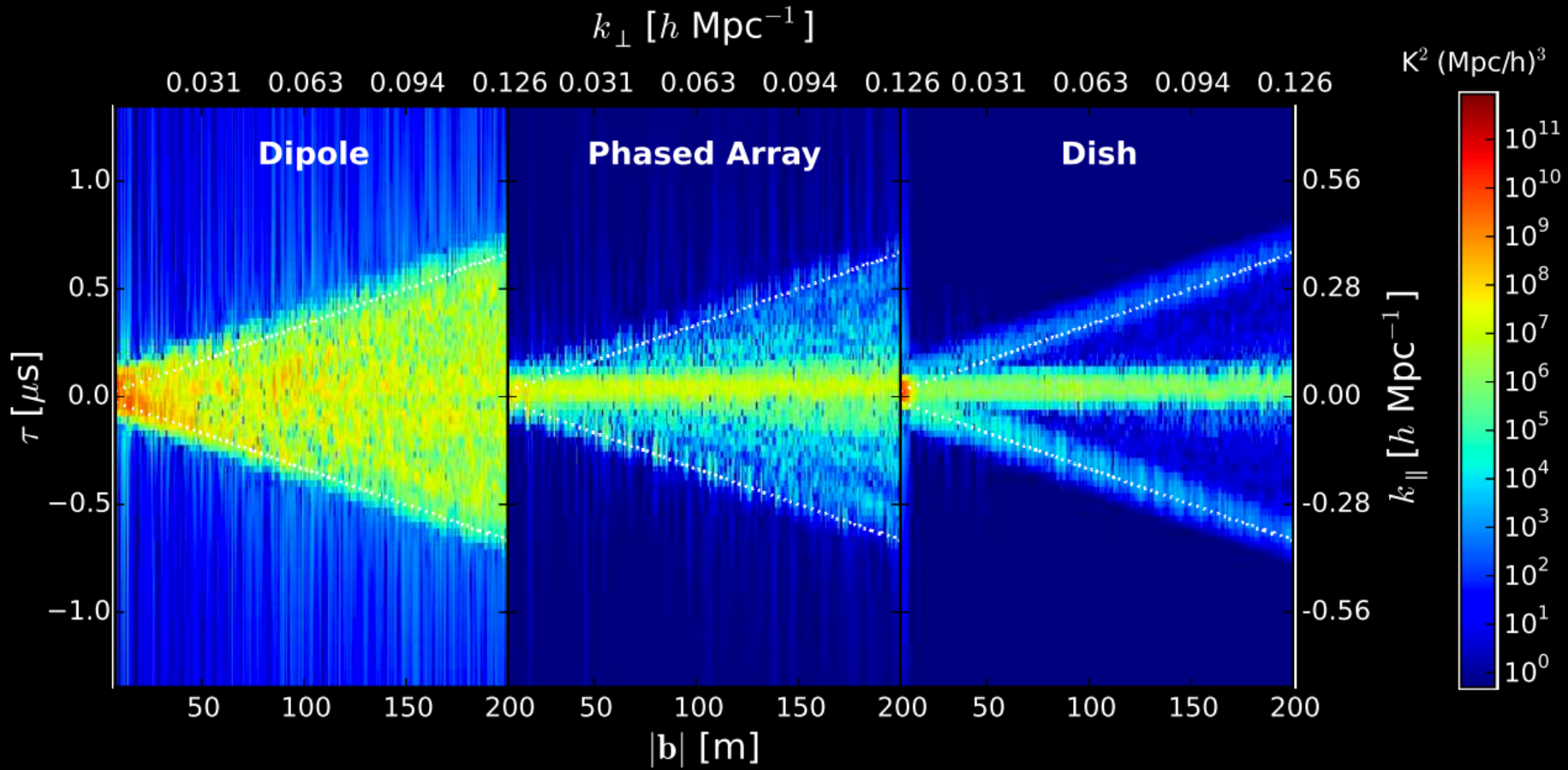
- Line Emission
  - Optically thin
  - Observable even into Dark Ages
  - Directly probes redshift evolution
- Sensitive to
  - Ly- $\alpha$ : Wouthuysen Field Effect
  - X-ray: Heating
  - UV: Ionization
- Integrated measure of galactic properties
  - includes low-mass tails
  - bubble size breaks some degeneracies
- Complementary to
  - CMB
  - optical probes of galaxies
  - other potential lines (e.g. CO, CII)

# The (21cm) Frequency Axis

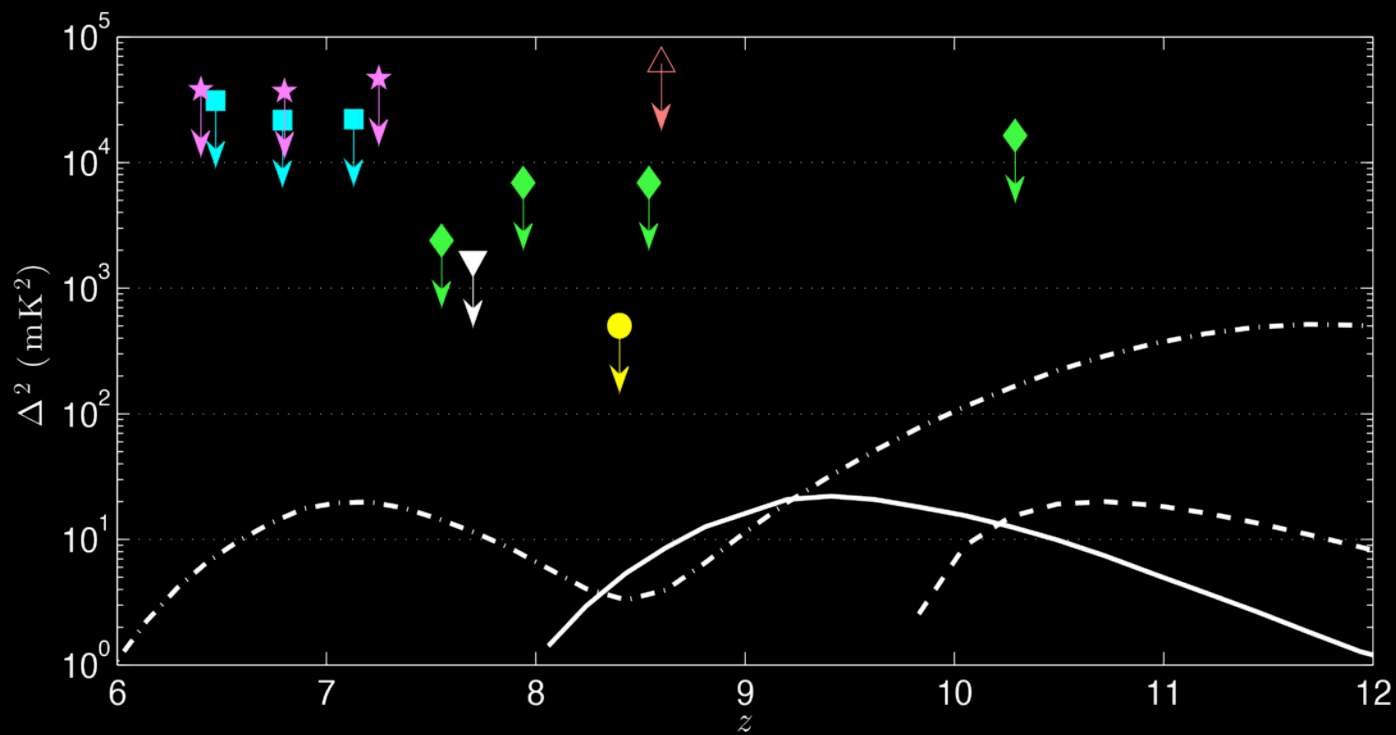
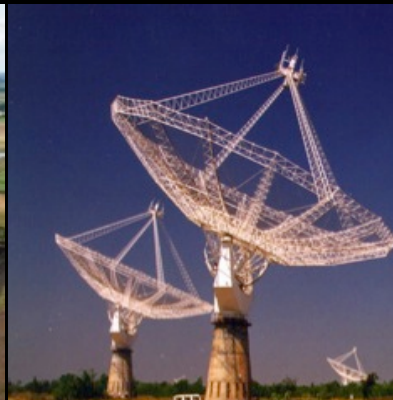
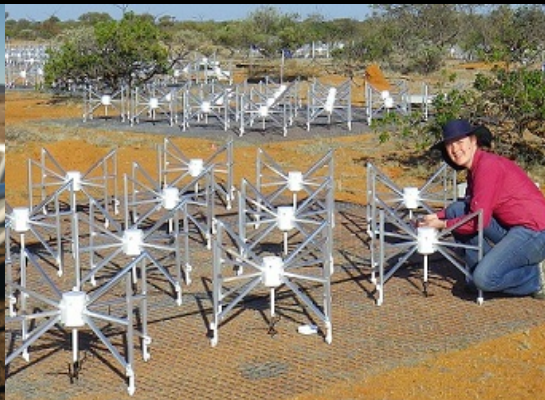
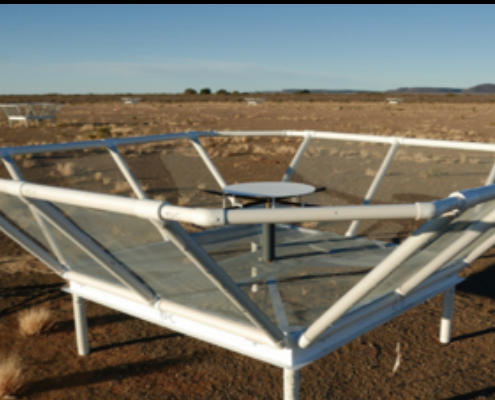








(Thyagarajan et al. 2015)

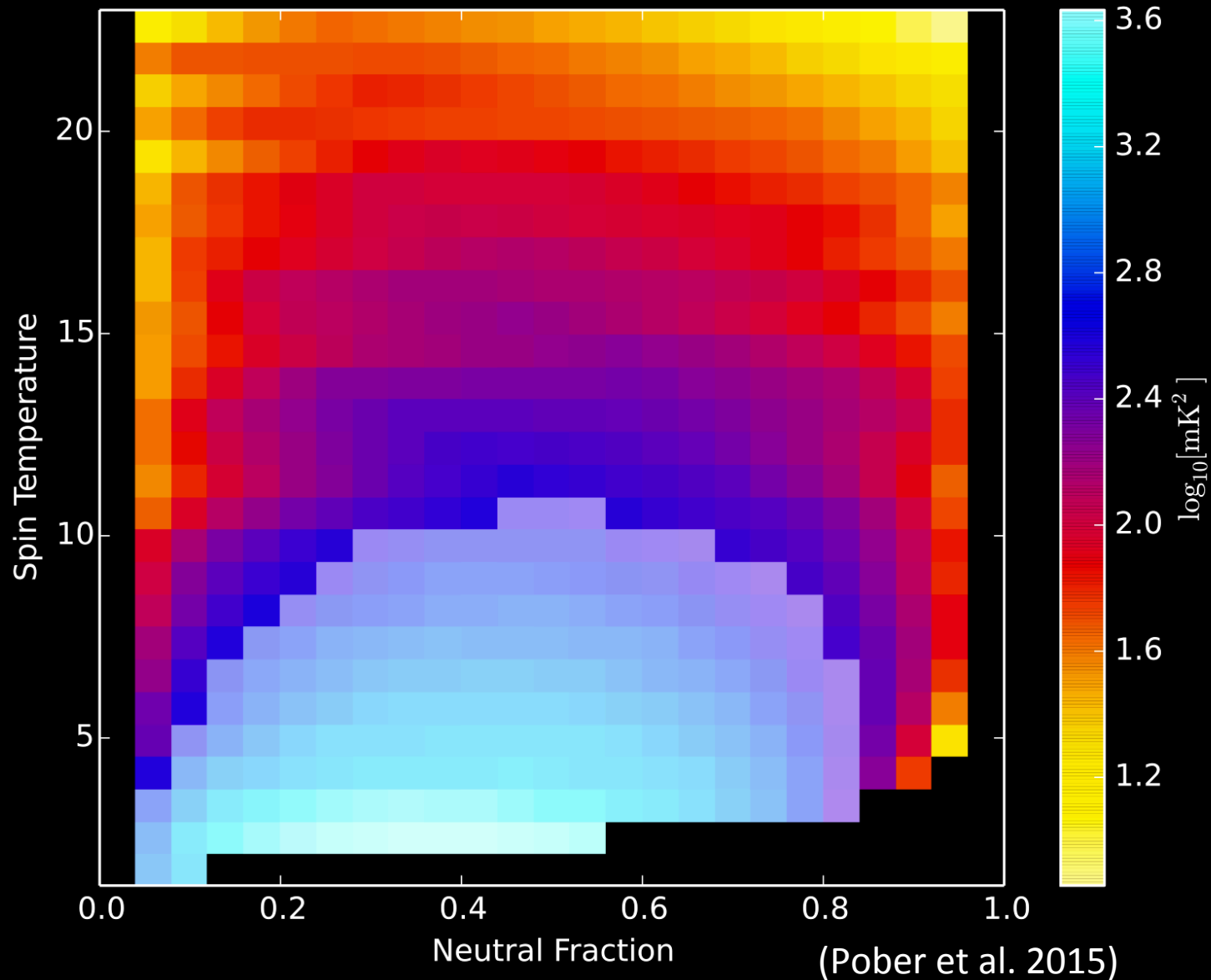


- |                   |                 |                    |
|-------------------|-----------------|--------------------|
| ● Ali, 2015       | ◆ Jacobs, 2015  | --- CDM Extreme    |
| ■ Beardsley, 2015 | ▼ Parsons, 2014 | — CDM Fiducial     |
| ★ Dillon, 2015    | △ Paciga, 2013  | -·-·- WDM Fiducial |



$$\delta T_b(\nu) \approx 9x_{\text{HI}}(1+\delta)(1+z)^{\frac{1}{2}} \left[ 1 - \frac{T_{\text{CMB}}(z)}{T_S} \right] \left[ \frac{H(z)/(1+z)}{dv_{\parallel}/dr_{\parallel}} \right] \text{ mK}$$

$k = 0.15$



**Location:** S30° 34', E21° 25' E (South Africa)

**Configuration:** 331 hex-pack, 21 outriggers

- **Min baseline:** 14.6m (7.8° scale)

- **Max baseline:** 1066m (9' beam)

**Array core:** 310m diameter

**Element:** 14m diameter (9° fov @150 MHz)

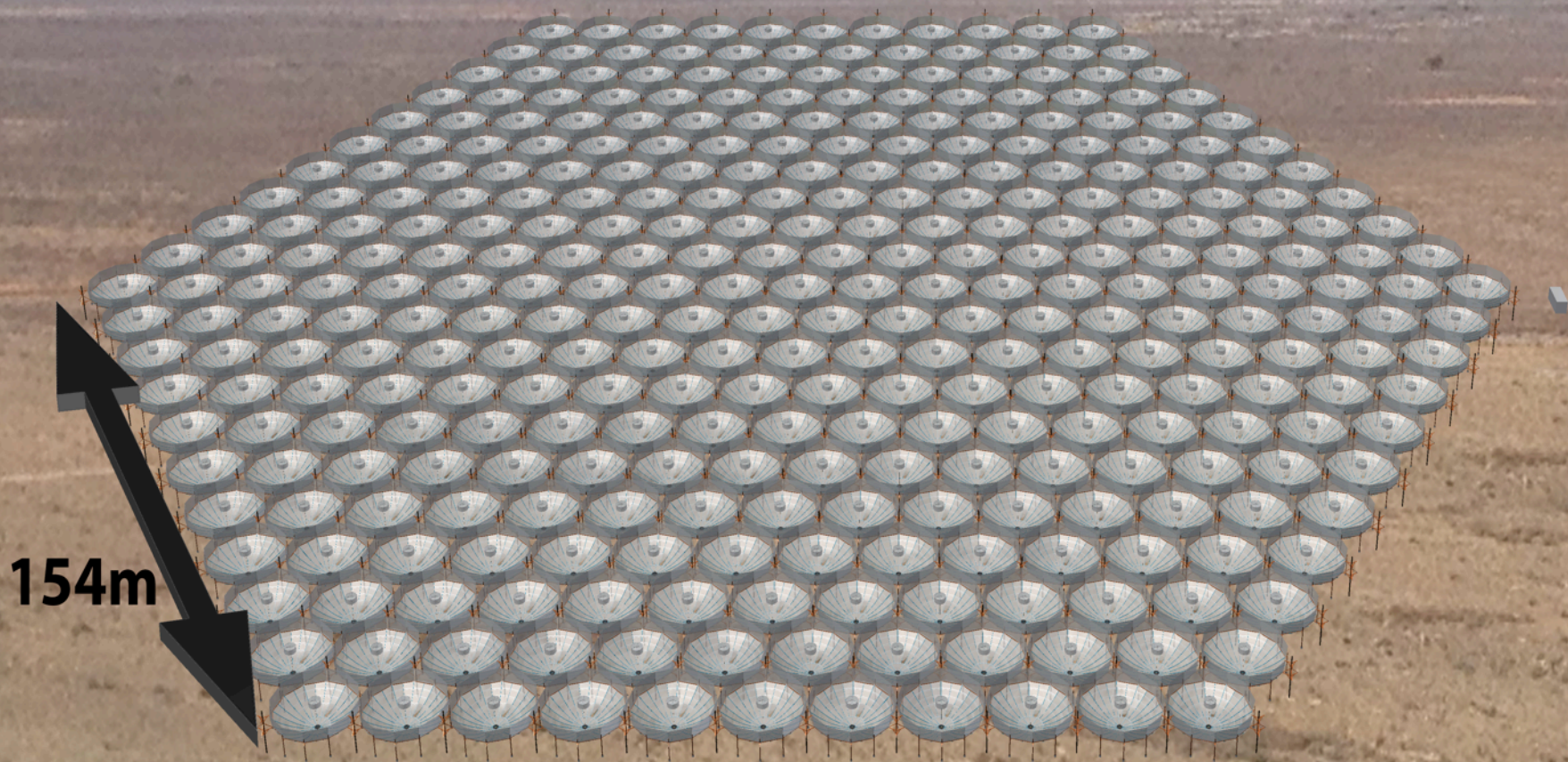
**Frequency**

- **Digitized:** 50 - 250 MHz

- **EOR band:** 100 - 200 MHz

- **Channel:** 97.7 kHz

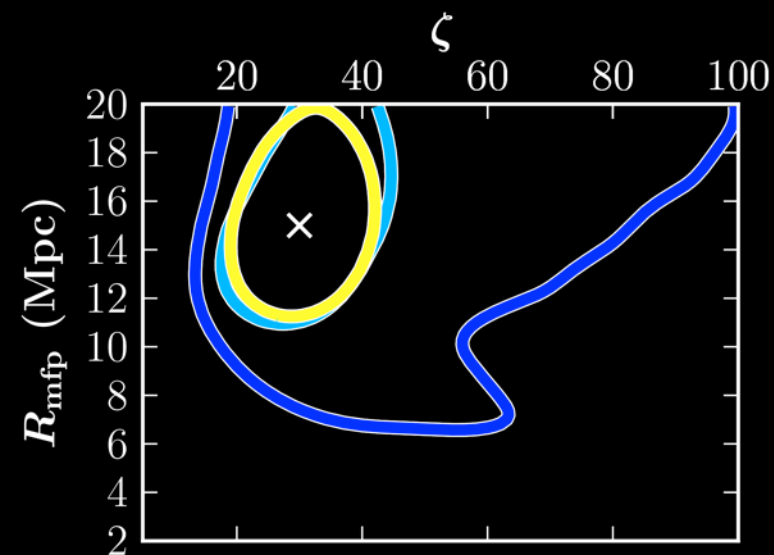
$$T_{\text{sys}} = 100 + T_{\text{sky}}$$



# EOR Power Spectrum - Array Sensitivities\*

Instrument	Avoidance		Subtraction	
	Drift	Track	Drift	Track
PAPER 128	1.56	-	4.46	-
MWA 128	0.66	0.86	2.50	3.15
LOFAR (core)	0.70	1.90	7.48	12.22
HERA 37	5.67	-	15.46	-
HERA 331	38.75	-	111.69	-
MWA 256	2.40	2.81	8.28	9.64
SKA1 Low	21.23	26.92	139.07	115.13

\* **Updated 21 cm Experiment Sensitivities**, J. Pober, HERA Memo #3,  
Table 2 (<http://reionization.org/memos>)



Instrument

Parameter (% error)

(multi- $z$ )

$\frac{\zeta}{\zeta_{\text{fid}}}$

$\frac{R_{\text{mfp}}}{R_{\text{fid,mfp}}}$

$\frac{\log_{10}(T_{\text{vir}}^{\text{min}})}{\log_{10}(T_{\text{vir, fid}}^{\text{min}})}$

LOFAR

1.32 (40.38)

1.03 (20.06)

1.05 (5.43)

HERA

1.03 (11.81)

1.00 (11.99)

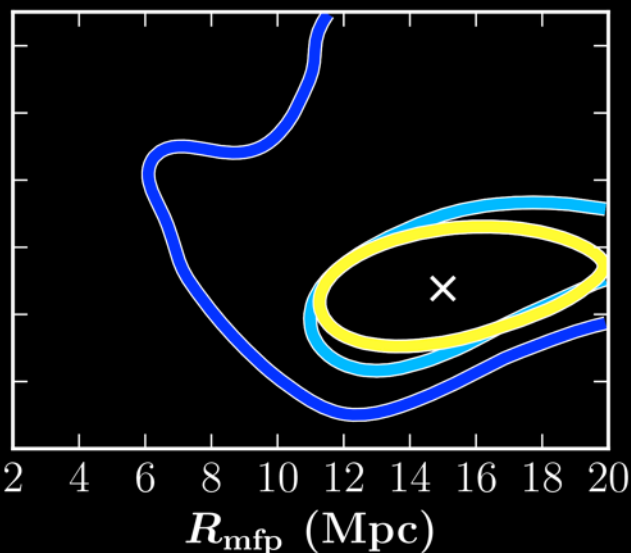
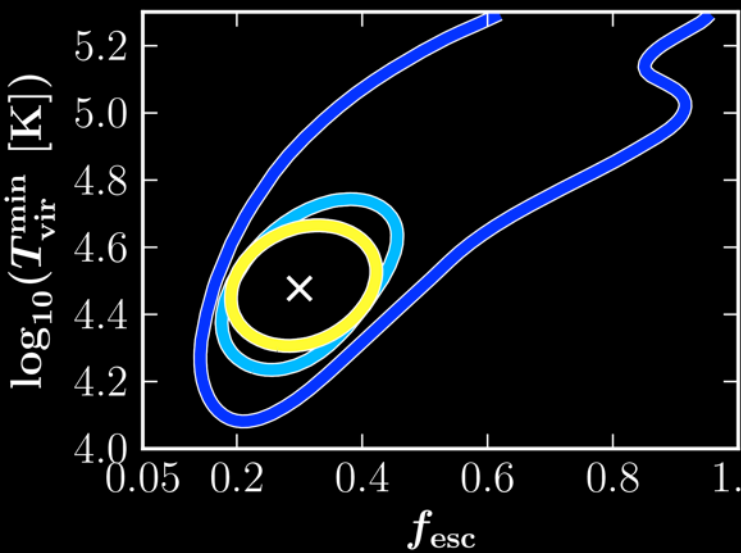
1.00 (1.95)

SKA

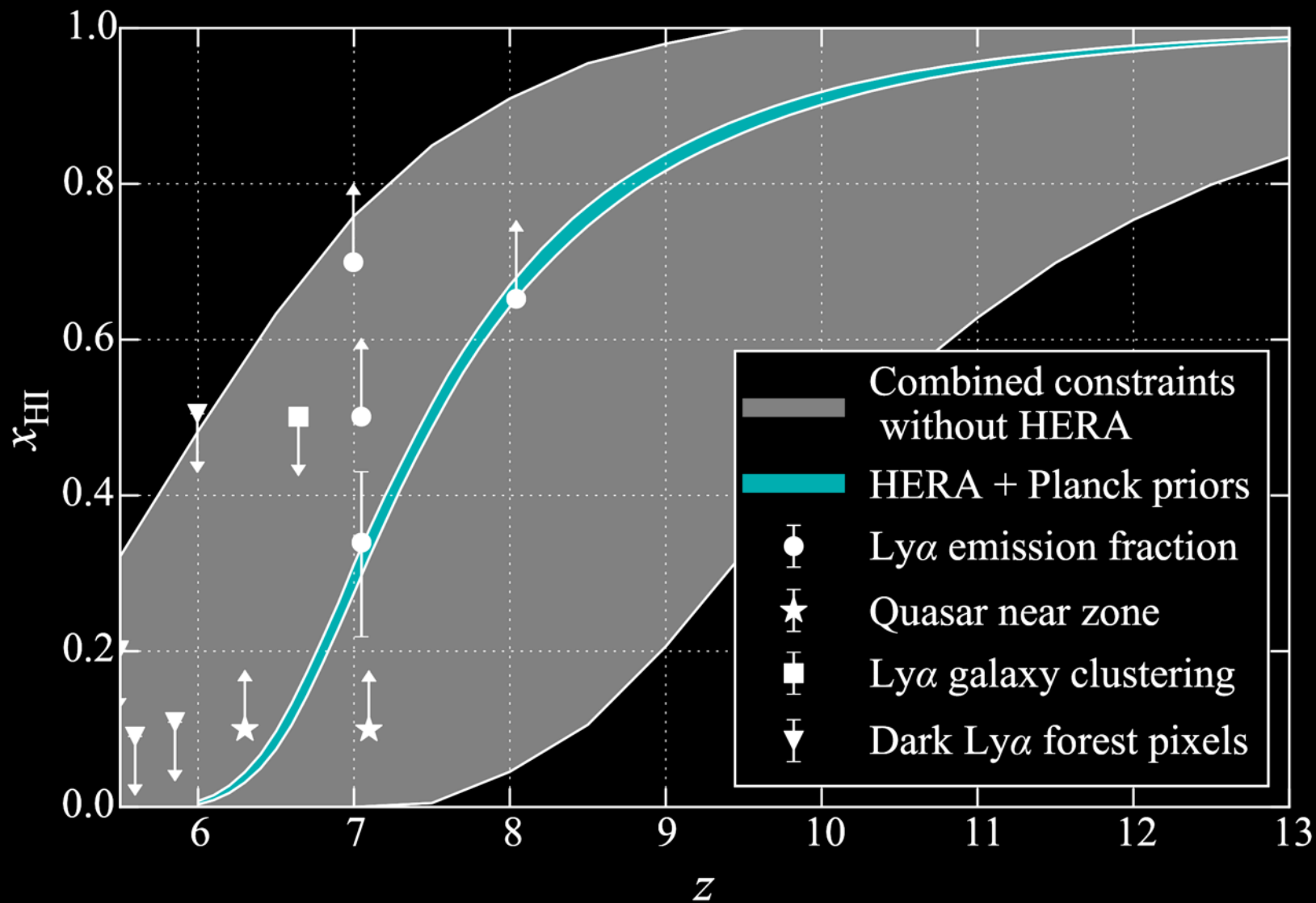
1.02 (6.11)

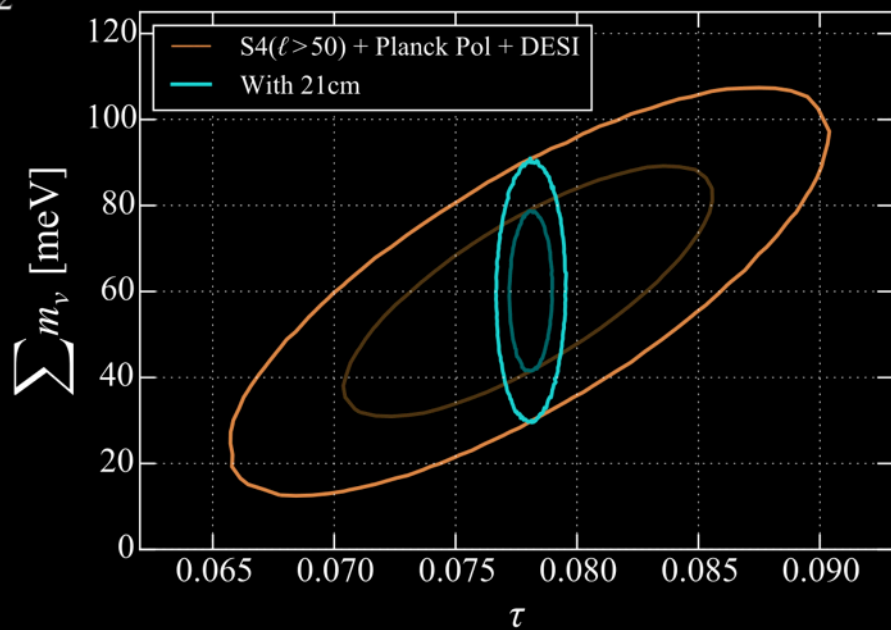
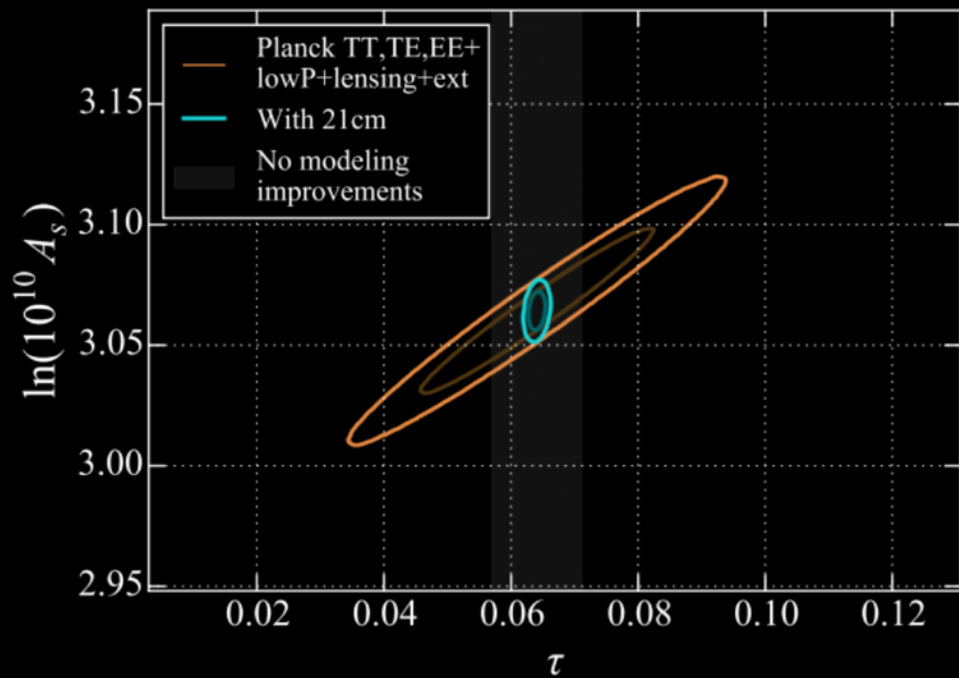
1.00 (10.04)

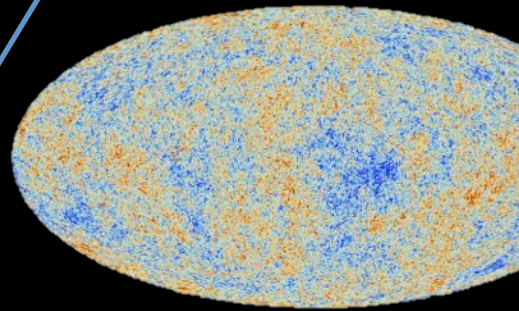
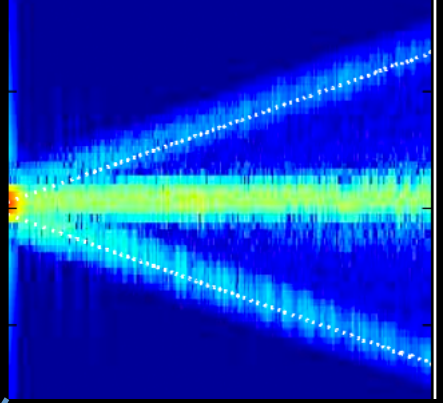
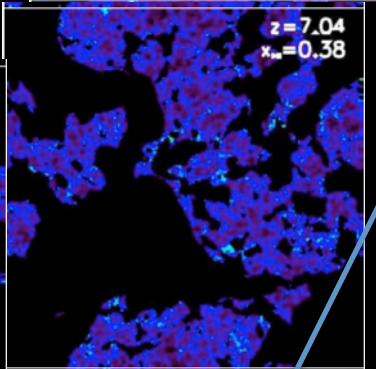
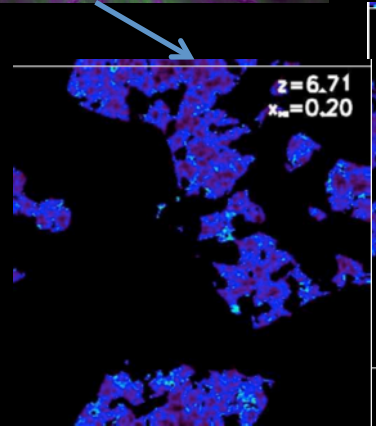
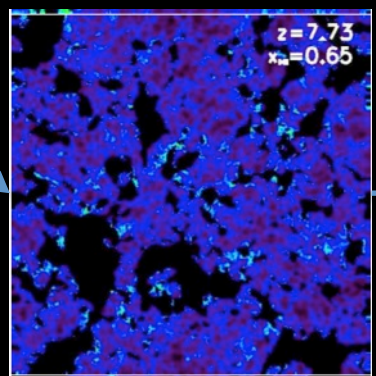
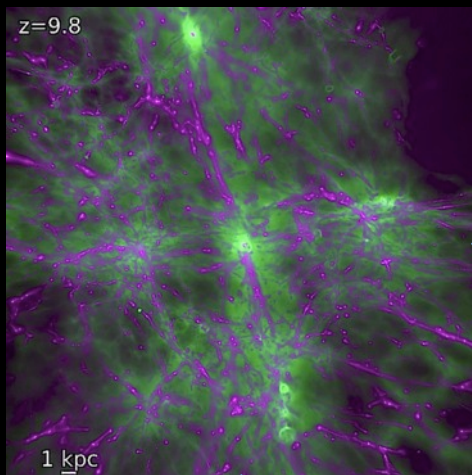
1.00 (0.96)



— LOFAR  $2\sigma$   
 — HERA  $2\sigma$   
 — SKA  $2\sigma$







$T_{\text{vir}} [\text{K}]$   
 $R_{\text{mfp}} [\text{Mpc}]$   
 $\zeta$

$\Omega_b h^2$   
 $\Omega_c h^2$   
 $100\theta_{\text{MC}}$   
 $\ln(10^{10} A_s)$   
 $n_s$   
 $\tau$

$$\rho_{\text{SFR}}(z) = a_p \frac{(1+z)^{b_p}}{1 + [(1+z)/c_p]^{d_p}}$$

