

Invited Talk
International Conference ‘Nonlinear Optics: East-West Reunion’ (NLO50 2011)
Suzdal, Russia
Friday 23 September, 2011

Nonlinear Optics of Vacuum

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Acknowledgments for Collaboration and advice: G. Mourou, F. Krausz, E. Goulielmakis, W. Leemans, K. Nakajima, K. Homma, D. Habs, P. Chomaz, H. Videau, T. Esirkepov, S. Bulanov, M. Kando, W. Sandner, A. Suzuki, M. Teshima, R. Assmann, R. Heuer, S. Karsch, F. Gruener, W. Chou, F. Takasaki, M. Nozaki, A. Chao, P. Bolton, J.P. Koutchouk, K. Ueda, Y. Kato, X. Q. Yan, R. Li, A. Ringwald, H. Ruhl, T. Ostermayr, S. Petrovic, C. Klier, B. Altschul, Y. K. Kim, M. Spiro, A. Seryi, A. Sergeev, A. Livak, K. Iqbal, C. Robilliard, J. Taran

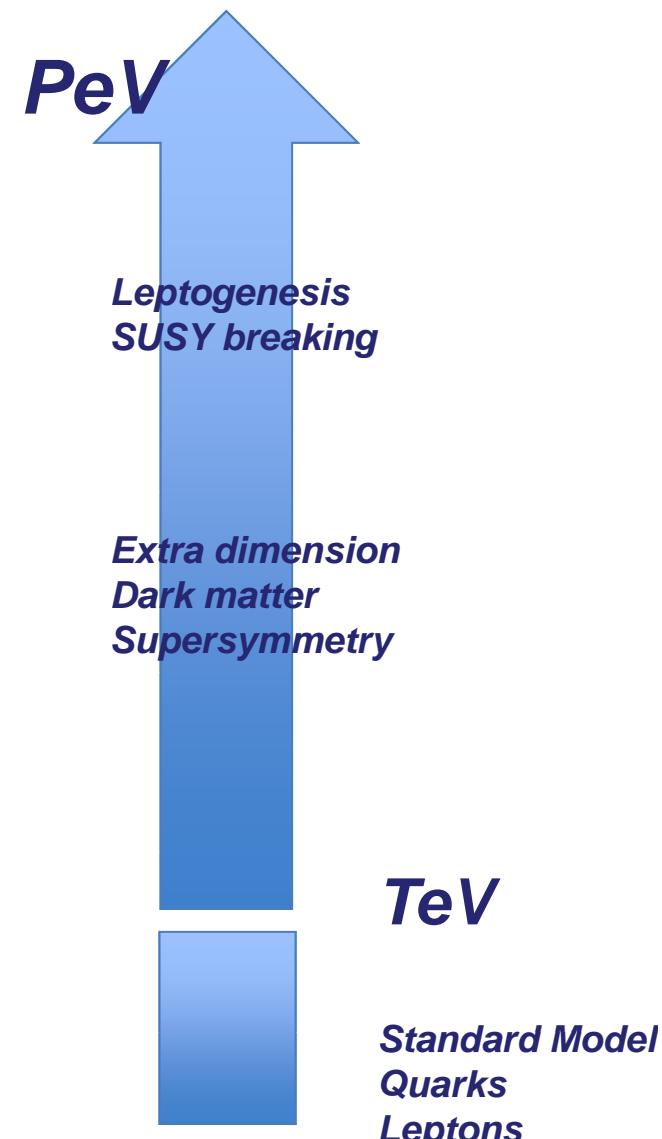
1. Suzuki's challenge in high energy physics
 - High energy frontier: TeV and beyond
 - A collider?
2. Non-collider paradigm
 - Vacuum texture and synchrotron radiation in high energy
 - Energy frontier at PeV with attosecond metrology without luminosity
3. **High Field** explores low energy new fields:
 - high field of **laser** (cf. high momentum)
 - Dark matter and dark energy fields in vacuum
 - 2nd harmonic, degenerate 4 wave mixing
4. zs streaking of vacuum by **laser** and γ photon
5. New initiative : **IZEST** = LIL compression, **XCELS** in Russia, etc.

IZEST's Mission: Responding to Suzuki's Challenge



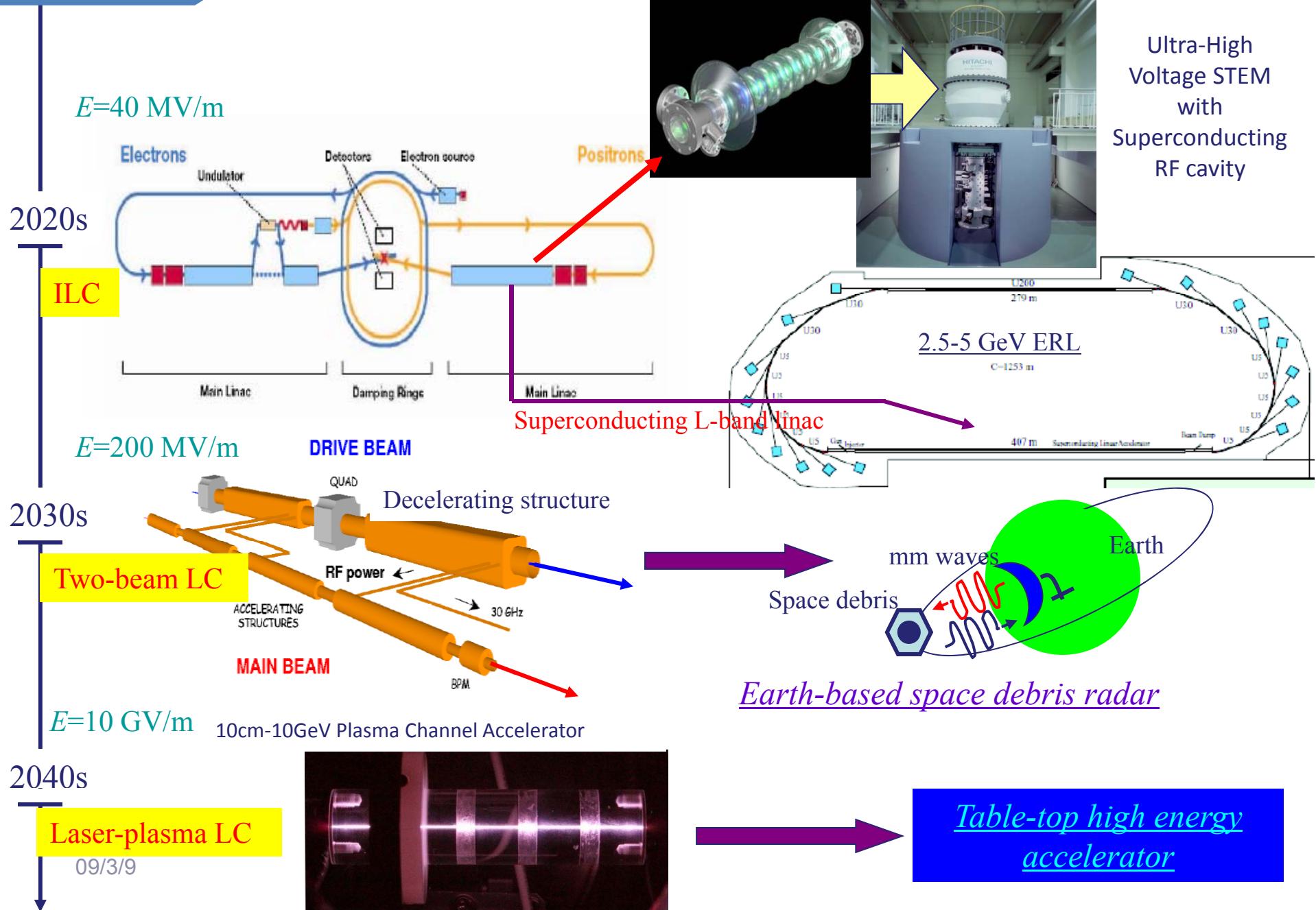
*Atsuto Suzuki:
KEK Director General,
ICFA Chair*

New Paradigm



Accelerator

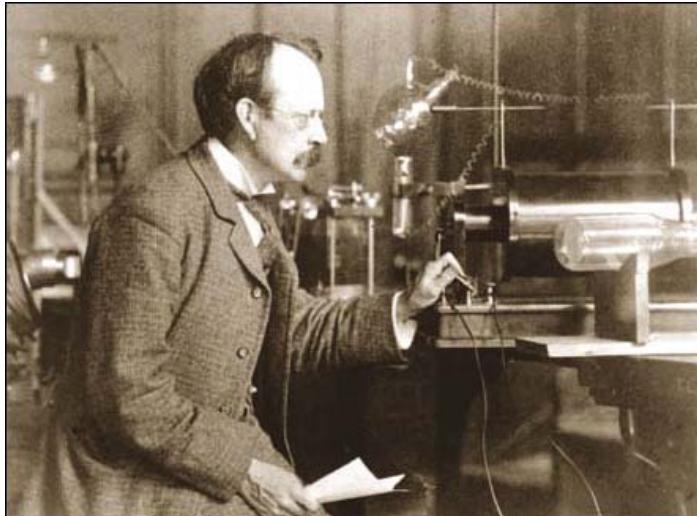
Evolution of Accelerators and their Possibilities (Suzuki,2008)



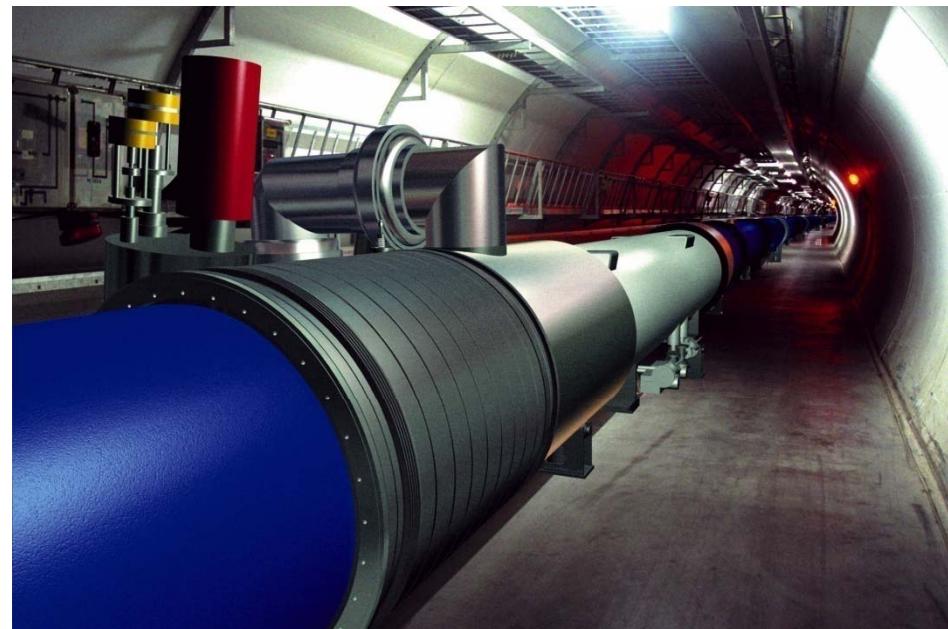


I ZEST

20th Century, the Electron Century Basic Research Dominated by Massive and Charged Particles



J. J. Thomson





IZEST

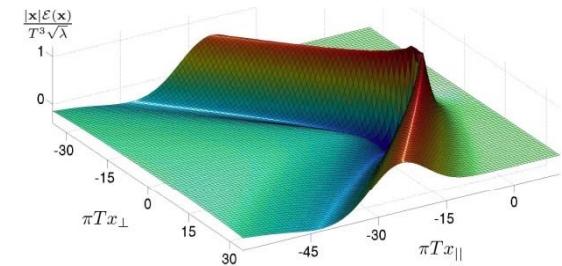
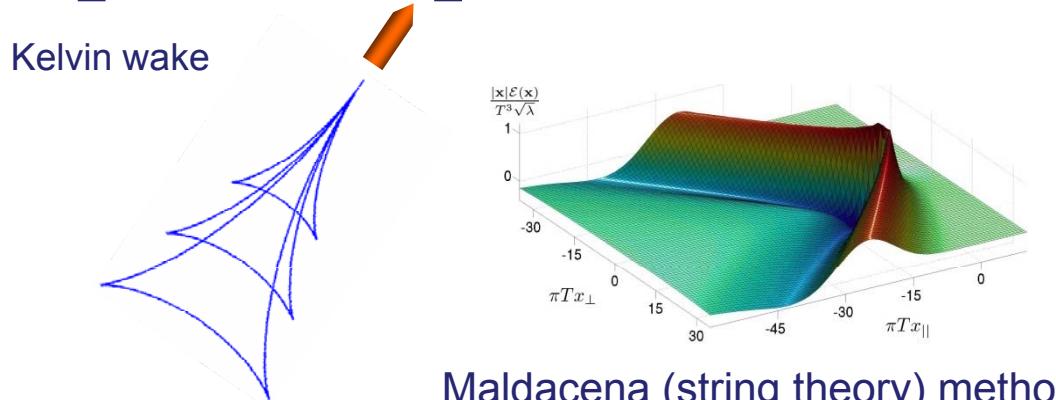
21st Century; the **Photon** Century

Could basic research be driven by the massless and chargeless particles; **Photons?**



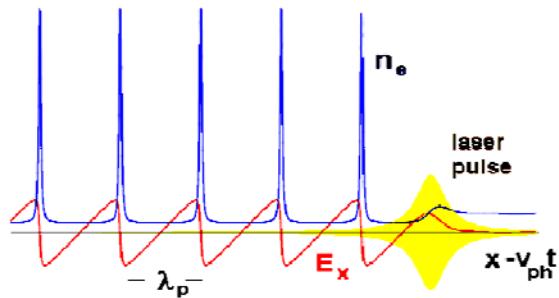
C. Townes

Laser Wakefield (LWFA): nonlinear optics in plasma



Maldacena (string theory) method:
QCD **wake** (Chesler/Yaffe 2008)

No wave breaks and wake **peaks** at $v \approx c$



← relativity
regularizes

(The density cusps.
Cusp singularity)

Wave **breaks** at $v < c$



Hokusai

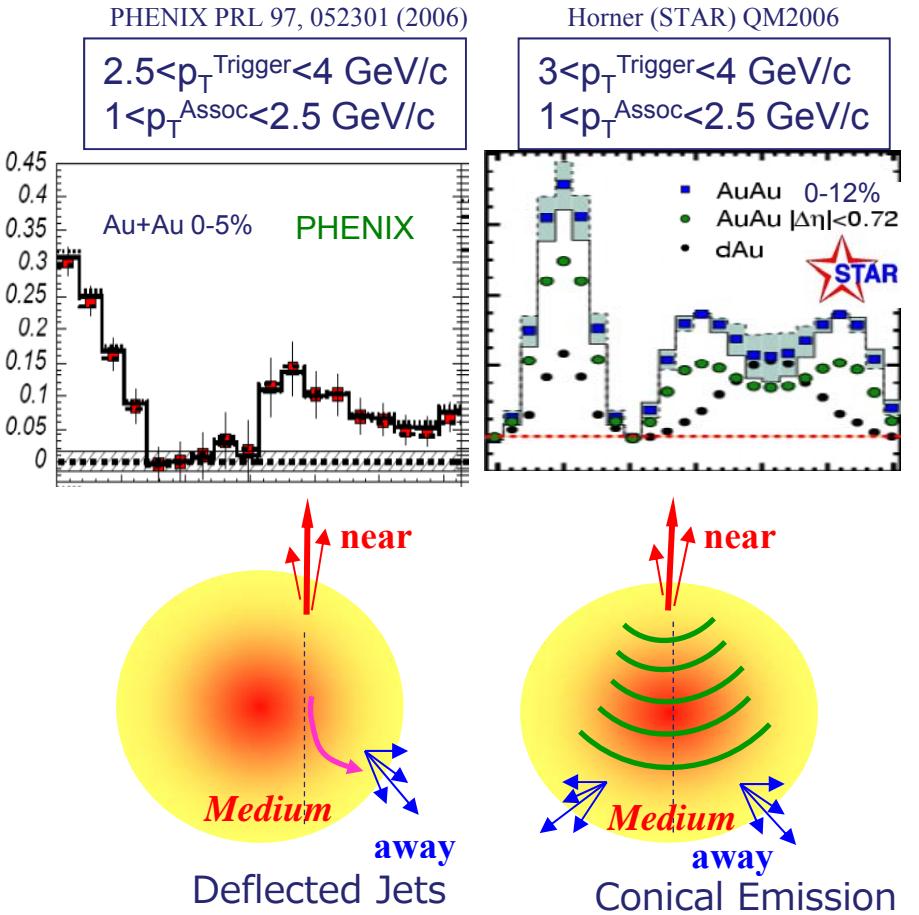


(Plasma physics vs.
String theory)



Nuclear Wake?

- BNL (and CERN) heavy ion collider: “**monojet**”
- Could be caused by:
 - Large angle gluon radiation (Vitev and Polsa and Salgado).
 - Deflected jets, due to flow (Armesto, Salgado and Wiedemann) and/or path length dependent energy loss (Chiu and Hwa).
 - Hydrodynamic conical flow from mach cone shock-waves (Stoecker, Casalderrey-Solana, Shuryak and Teaney, Renk, Ruppert and Muller).
 - Cerenkov gluon radiation (Dremin, Koch).
- Jet quenching: collective deceleration by wakefield?
 - LWFA method, or Maldacena method?



Density scalings of **LWFA** for collider

Accelerating field E_z	$\propto n_e^{1/2}$
Focusing constant K	$\propto n_e^{1/2}$
Stage length L_{stage}	$\propto n_e^{-3/2}$
Energy gain per stage W_{stage}	$\propto n_e^{-1}$
Number of stages N_{stage}	$\propto n_e$
Total linac length L_{total}	$\propto n_e^{-1/2}$
Number of particles per bunch N_b	$\propto n_e^{-1/2}$
Laser pulse duration τ_L	$\propto n_e^{-1/2}$
Laser peak power P_L	$\propto n_e^{-1}$
Laser energy per stage U_L	$\propto n_e^{-3/2}$
Radiation loss $\Delta\gamma$	$\propto n_e^{1/2}$
Radiative energy spread σ_γ/γ_f	$\propto n_e^{1/2}$
Initial normalized emittance ϵ_{n0}	$\propto n_e^{-1/2}$
Collision frequency f_c	$\propto n_e$
Beam power P_b	$\propto n_e^{1/2}$
Average laser power P_{avg}	$\propto n_e^{-1/2}$
Wall plug power P_{wall}	$\propto n_e^{1/2}$

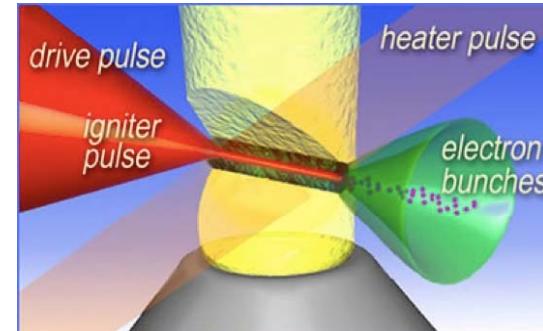
PeV Accelerator



*With conventional Technology
The accelerator would Girdle the Earth:
Fermi's vision (1954)*

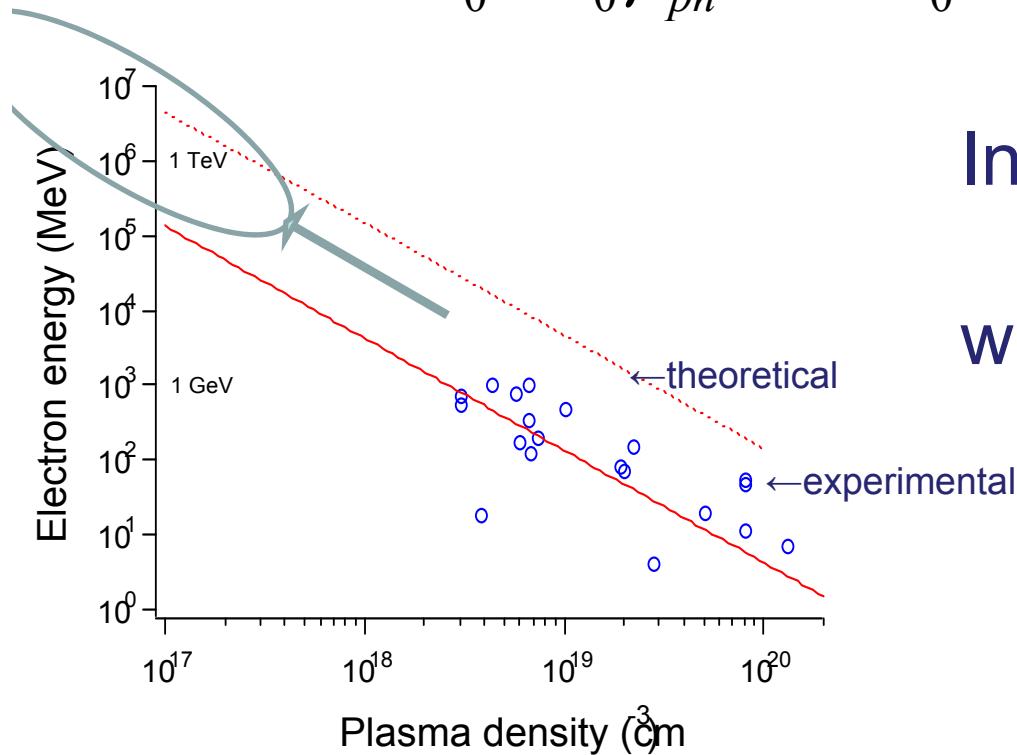


*1km **laser** plasma accelerator
with **LIL** or **LMJ**
(Vision 2011)*



Theory of wakefield toward extreme energy

$$\Delta E \approx 2m_0c^2a_0^2\gamma_{ph}^2 = 2m_0c^2a_0^2\left(\frac{n_{cr}}{n_e}\right), \text{ (when 1D theory applies)}$$



$$L_d = \frac{2}{\pi} \lambda_p a_0^2 \left(\frac{n_{cr}}{n_e} \right), \quad L_p = \frac{1}{3\pi} \lambda_p a_0 \left(\frac{n_{cr}}{n_e} \right),$$

dephasing length pump depletion length

In order to avoid wavebreak,
 $a_0 < \gamma_{ph}^{1/2}$,
where

$$\gamma_{ph} = (n_{cr}/n_e)^{1/2}$$

Adopt:
LMJ laser (3MJ)
 \rightarrow 0.7PeV
(with Kando, Teshima)

γ -ray signal from primordial GRB

LETTERS

NATURE

(Abdo, et al, 2009)

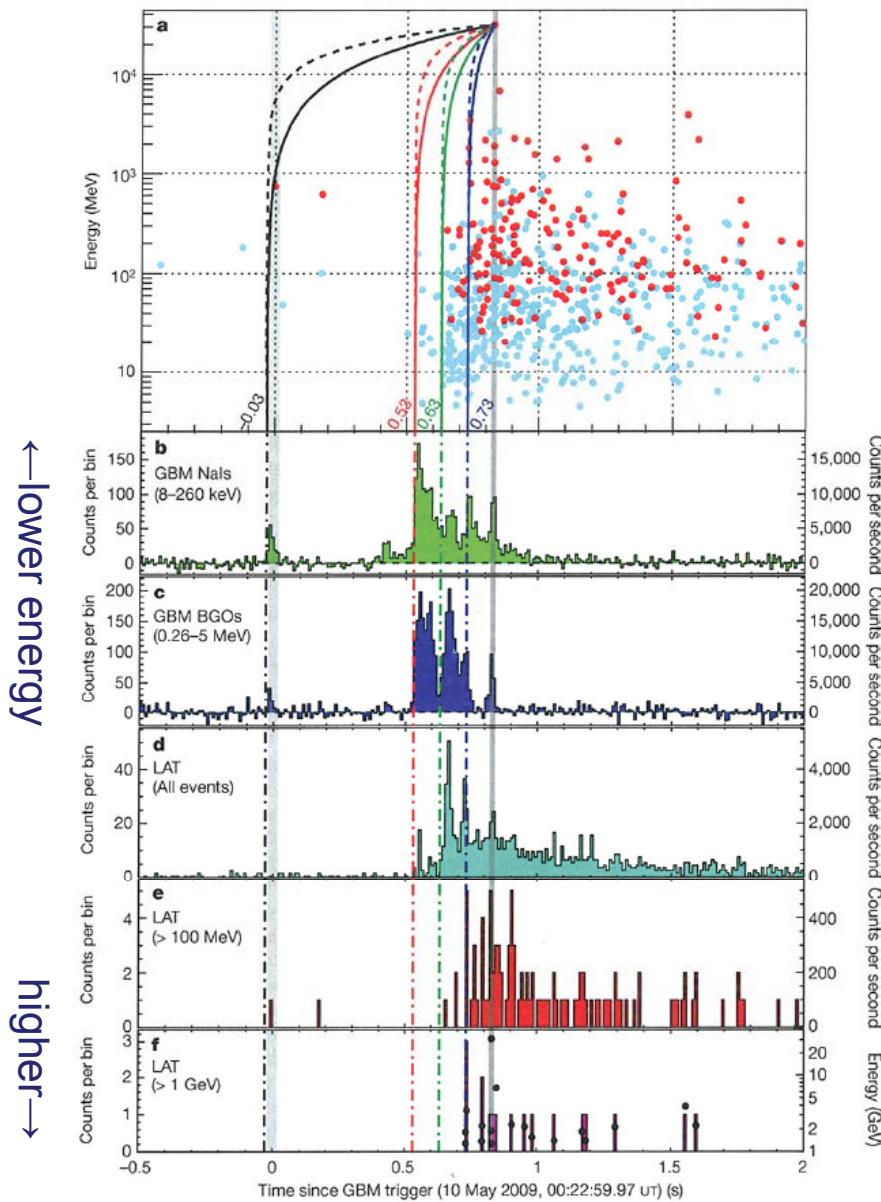


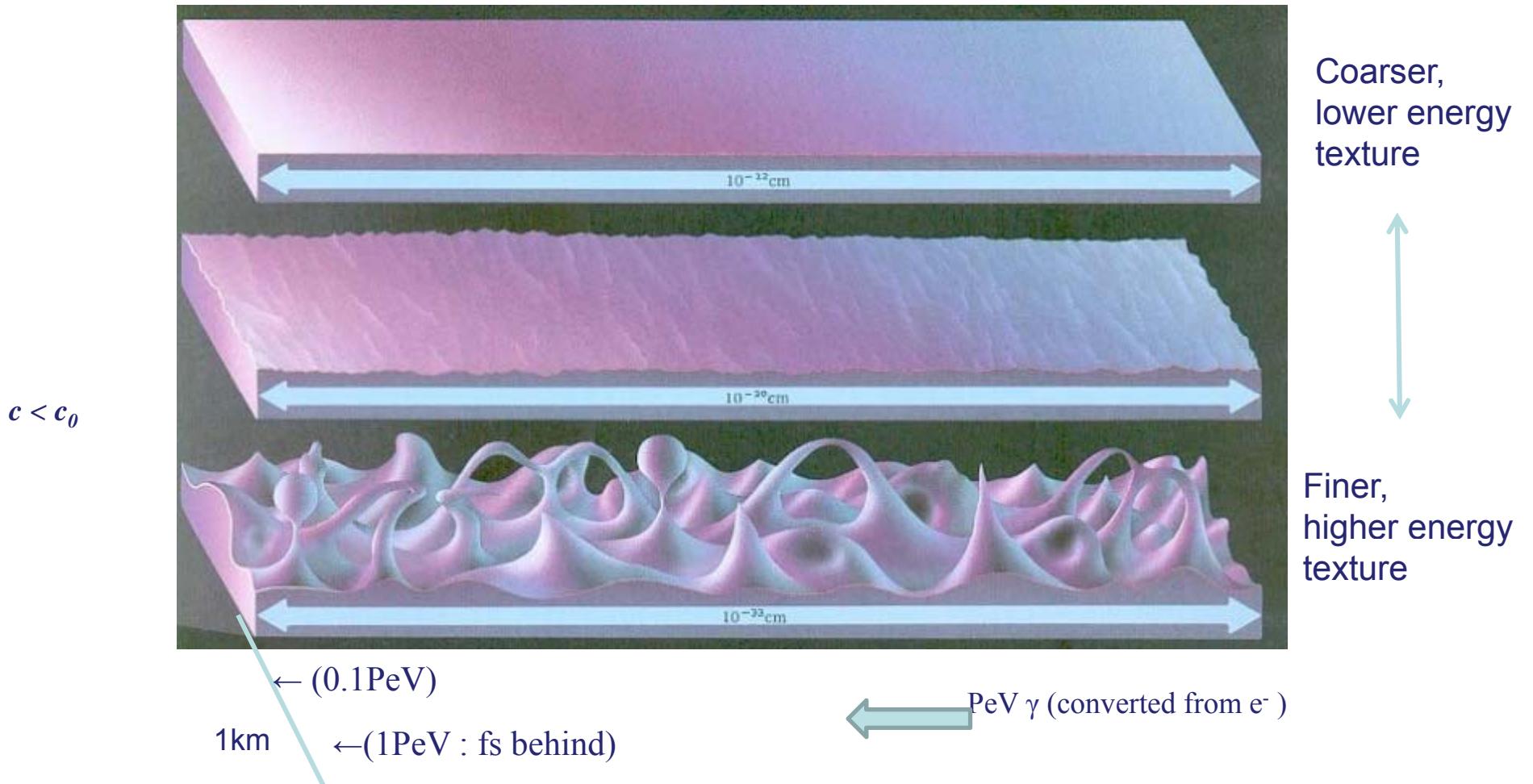
Figure 1 | Light curves of GRB 090510 at different energies. a, Energy lowest to highest energies. f also overlays energy versus arrival time for each

*Energy-dependent photon speed ?
Observation of primordial Gamma Ray Bursts (GRB)
(limit is pushed up close to Planck mass)*

**Lab PeV γ (from e-)
can explore this
with control**

Feel vacuum texture: PeV energy γ

Laser acceleration → controlled laboratory test to see quantum gravity texture on photon propagation (Special Theory of Relativity: c_0)



Extreme High Energy and Synchrotron Radiation

$E > 30\text{TeV}$: untested territory for Lorentz invariance

with a modified Lorentz factor

(B. Altschul, 2008)

$$\tilde{\gamma} = \frac{1}{\sqrt{1 + 2\delta_\gamma(\hat{v}) - v^2}}. \quad (13)$$

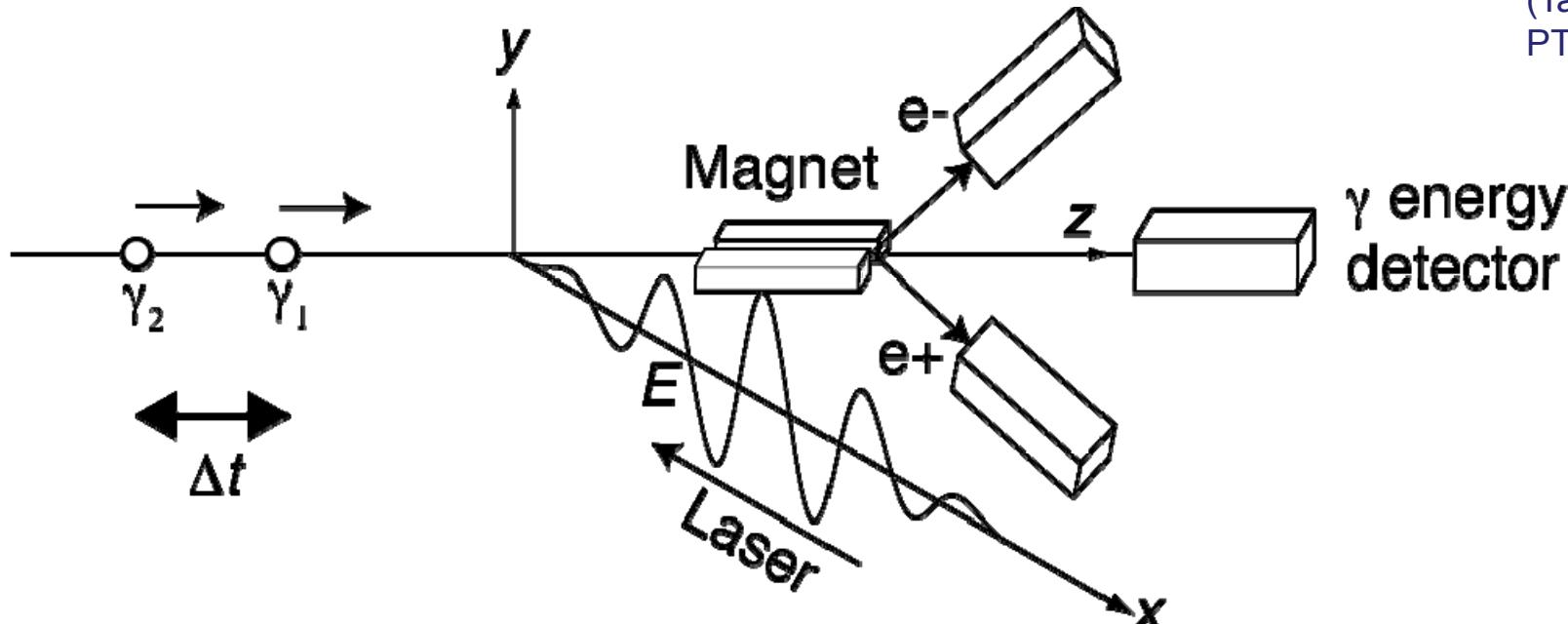
The power radiated would then be $P = \frac{e^2 a^2}{6\pi m^2} \tilde{\gamma}^4$.] For ultrarelativistic particles, $\gamma \approx [2(1 - v)]^{-1/2}$ increases very rapidly as a function of v , since $\frac{d\gamma}{dv} = v\gamma^3 \approx \gamma^3$. The modified expression for $\vec{v}(\vec{p})$ changes the radiated power $P(\vec{p})$ to

$$P(\vec{p}) = P_0(\vec{p}) \{1 + 4\gamma^2 [\delta(\hat{p}) - \delta_\gamma(\hat{p})]\}, \quad (14)$$

Synchrotron radiation
radiation

↑ Lorentz violating term ($>30\text{TeV}$)

Attosecond Metrology of PeV γ Arrivals



(Tajima, Kando,
PTP, 2011)

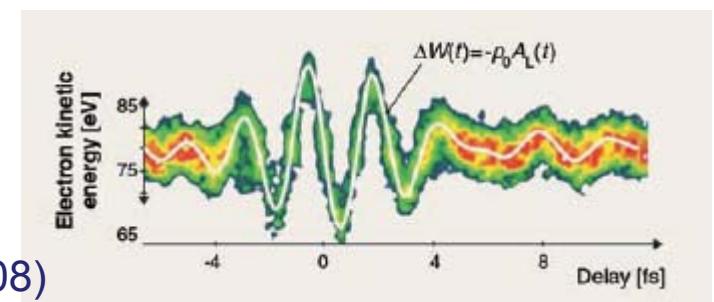
Narozhny, Nikisho, Ritus

High energy γ - induced Schwinger breakdown (Narozhny, 1968)

CEP phase sensitive electron-positron acceleration

Attosecond electron streaking

γ - energy tagging possible

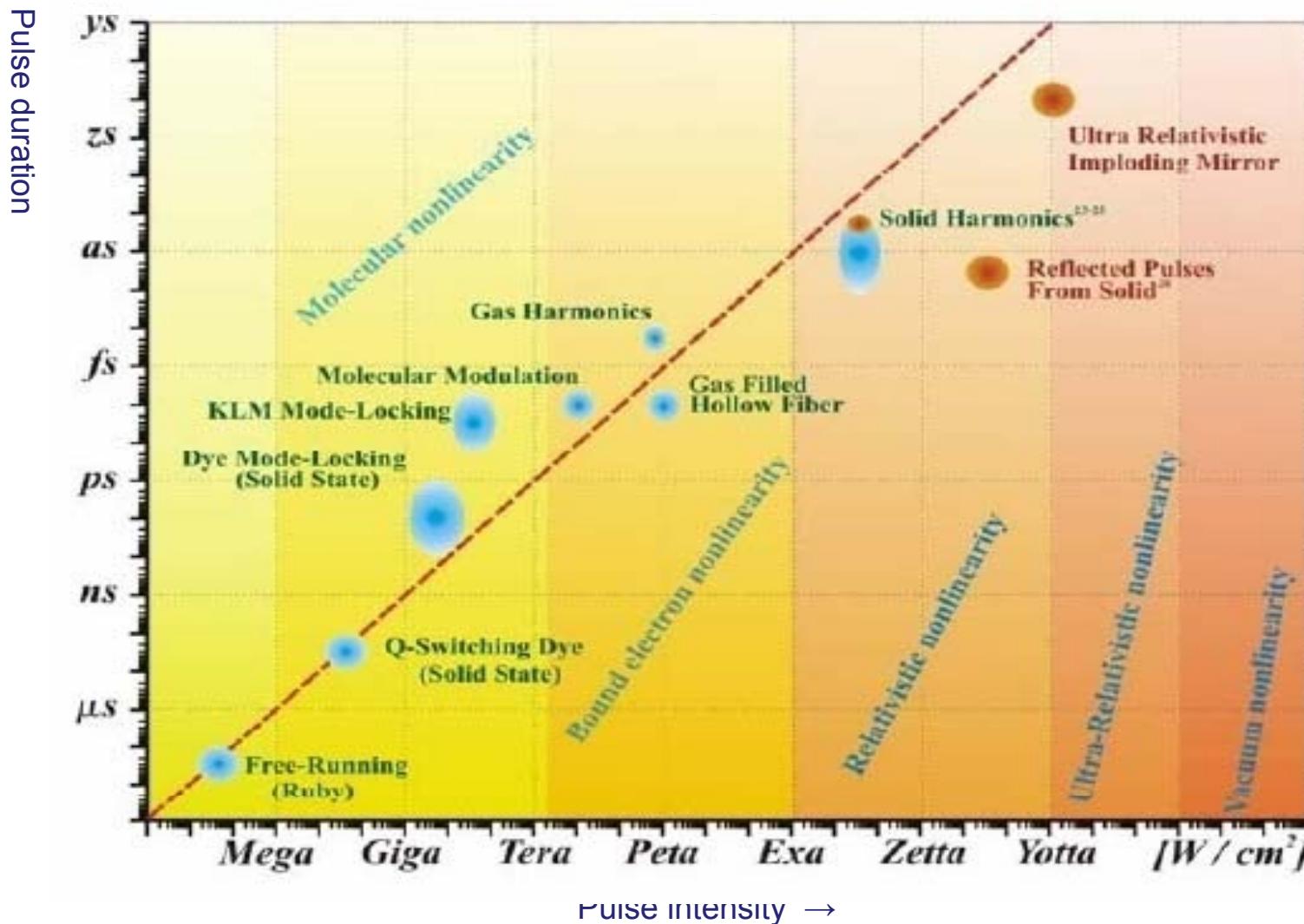


Goulielmakis(2008)

The Conjecture

(← physics: “Matter is **nonlinear**”)

“The more rigid nonlinearity, the more intense to manipulate it”;
rigidity vs. pulse length)



(Mourou / Tajima, science, 2011)

Streaking Vacuum

(from atomic physics to QED vacuum physics)

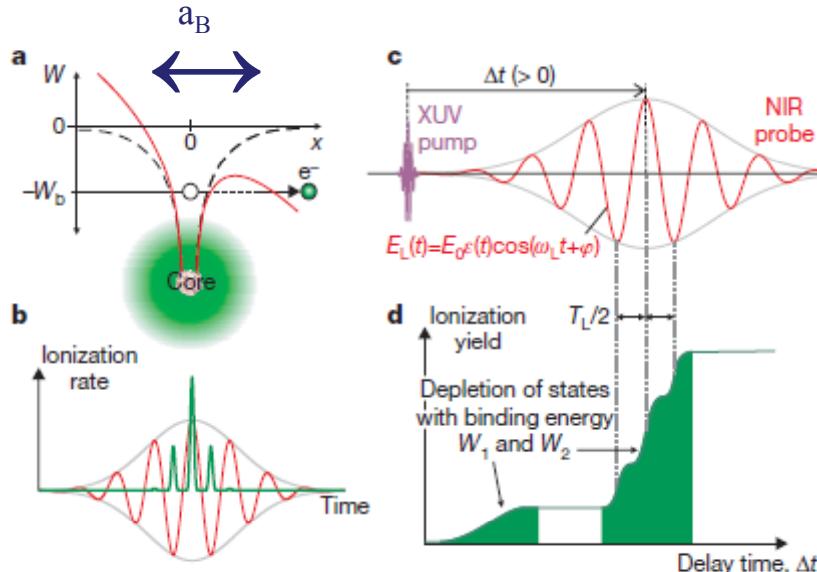
vacuum

Gamma photon 'ionization'

XUV streaking

→ zeptosecond dynamics

Uiberacker et al. (2007)



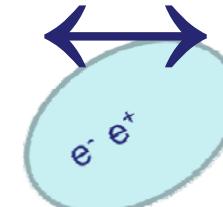
XUV photon ionization
Laser streaking
→ attosecond dynamics

atom

$$E_S/E_K = \alpha^{-3}; P_{c\ vac}/P_c = \alpha^{-6}$$

size

$$\lambda_C = \alpha a_B$$



depth of potential

$$\Phi = \alpha^{-2} W_B$$

$$R_{e^+e^-} \propto \exp\left(-\left(\frac{8}{3}\right)\left(\frac{m}{\omega}\right)\left(\frac{E_S}{E}\right)\right)$$

Nikishov(1964)

Nonperturbative:

$$W_{\parallel} = \frac{32\pi^2 m^2}{33\hbar^2} \left(\frac{x}{2\pi}\right)^{1/2} e^{-x/2}, \quad W_{\perp} = 2W_{\parallel}, \quad x \ll L. \quad (36')$$

For large values of x we essentially have $n \gg 1$ in the integrals (36). Using this fact, we obtain

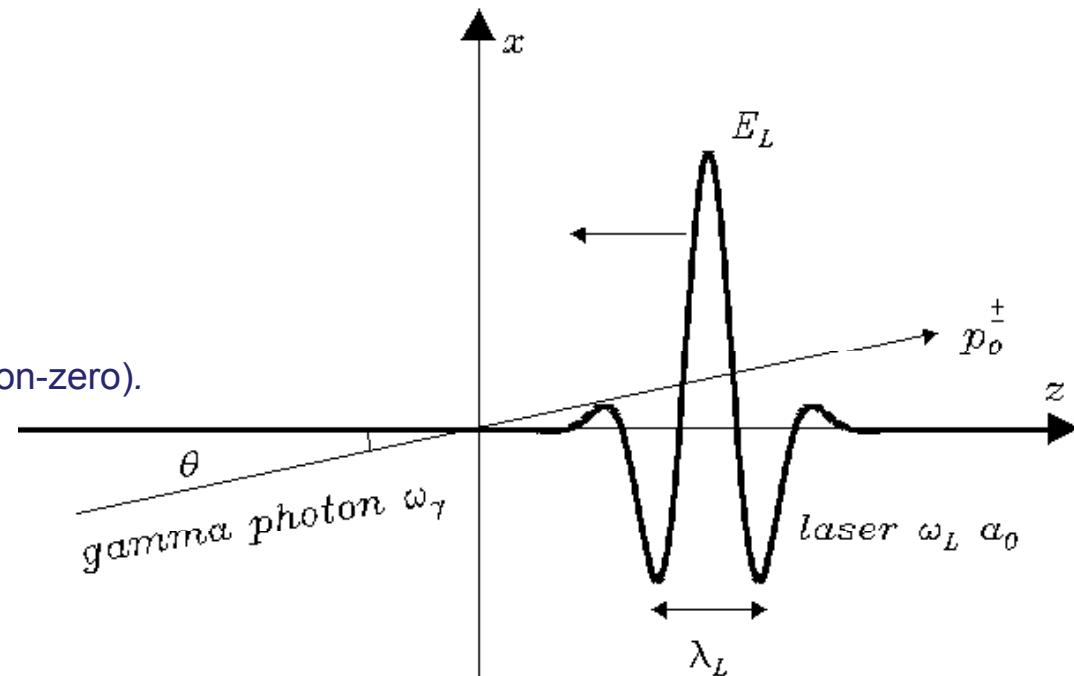
$$W_{\parallel} = \frac{32\Gamma^2(5/2)}{33\pi^2 n} \left(\frac{2n}{3}\right)^{1/2}, \quad W_{\perp} = \frac{2}{3} W_{\parallel}, \quad n \gg 1. \quad (36'')$$

γ -photon induced vacuum streaking by lasers

Schwinger-Nikishov amplitude

$$a_0^{SN} = (mc^2/\hbar\omega_L)(mc^2/\hbar\omega_\gamma)$$

(We need to make the Schwinger invariant non-zero).

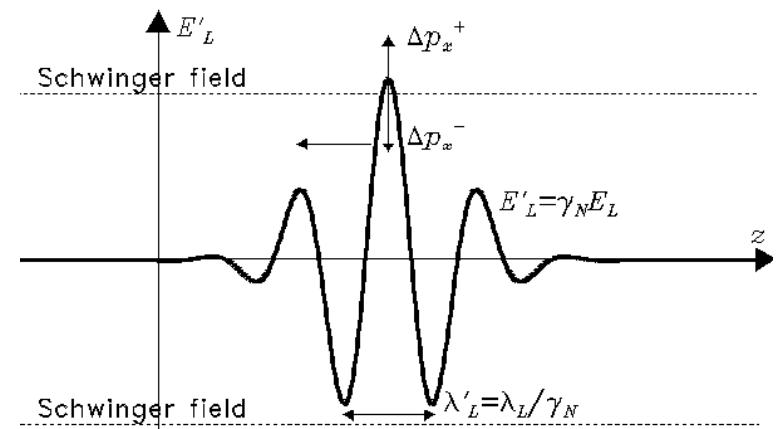


time resolution of streaking on the Nikishov frame (the Nikishov $\gamma_N = \hbar\omega_\gamma/mc^2$)

$$\Delta t' = [2(\hbar/mc^2)/(a_0^2 \omega_L'^2)]^{1/3}.$$

Necessary **laser** amplitude:

$$a_0^{\text{res}} = 2 (mc^2/\hbar\omega_L').$$

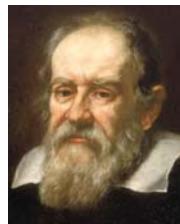


Laser fits the gaping hole

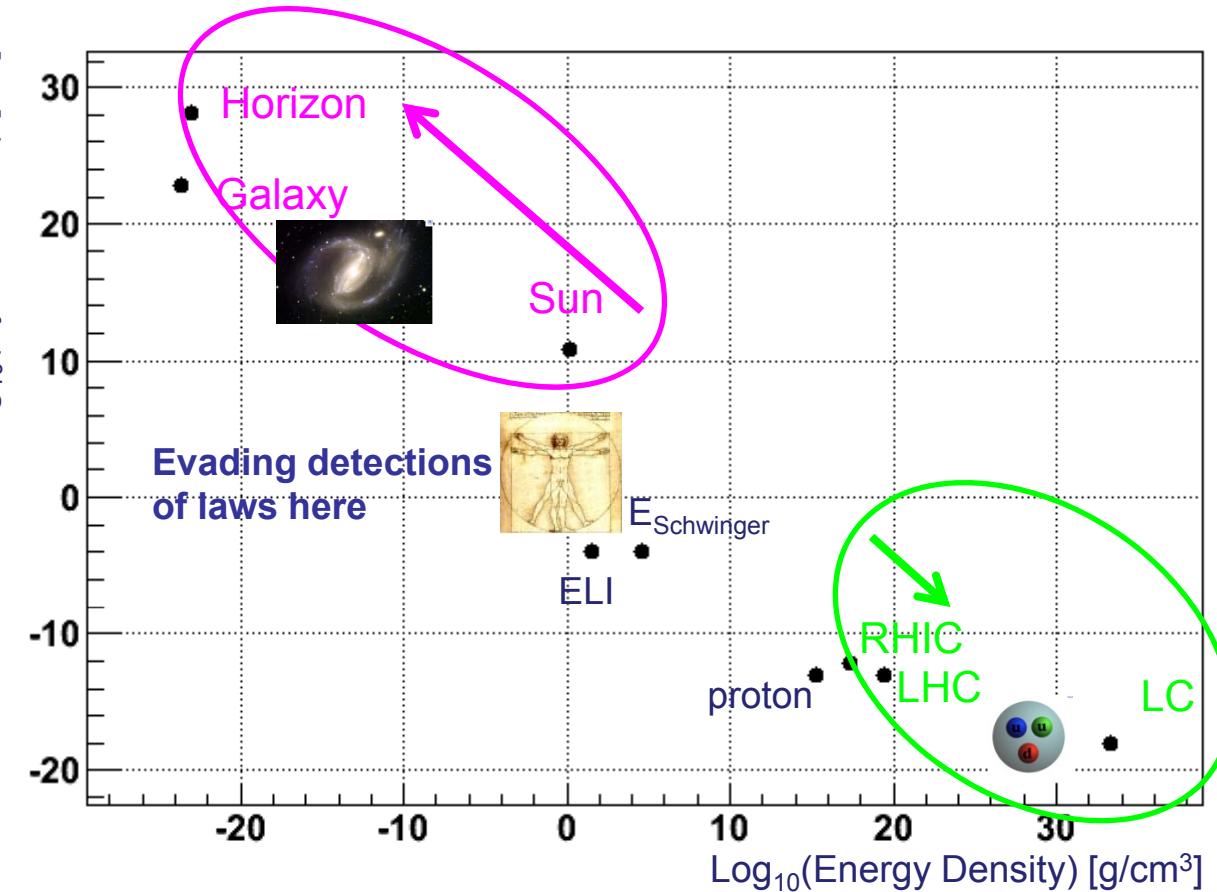
in search of unknown fields:
dark matter/dark energy



Cosmological
observation



Log₁₀(System Size) [cm]



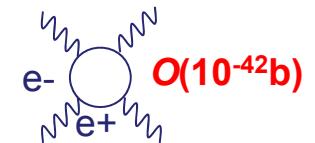
High energy
collider

Domains of physical laws

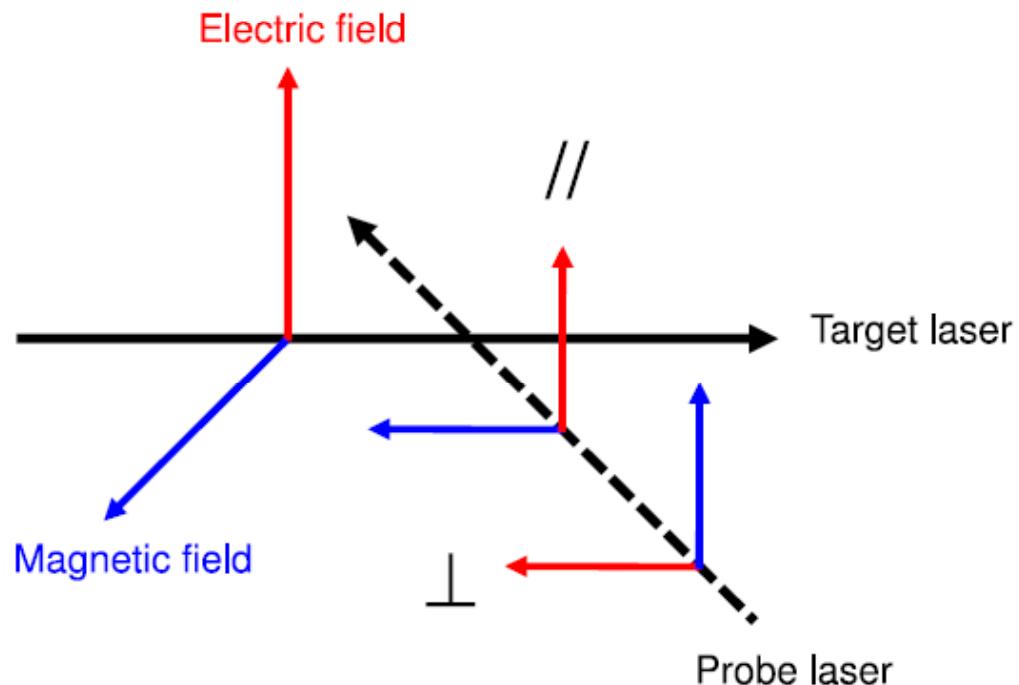
Birefringence by QED in eV range

Euler-Heisenberg one-loop Lagrangian

$$L_{QED} = \frac{1}{360} \frac{\alpha^2}{m^4} [4(F_{\mu\nu}F^{\mu\nu})^2 + 7(F_{\mu\nu}\tilde{F}^{\mu\nu})^2]$$



Refractive index depends on polarizations

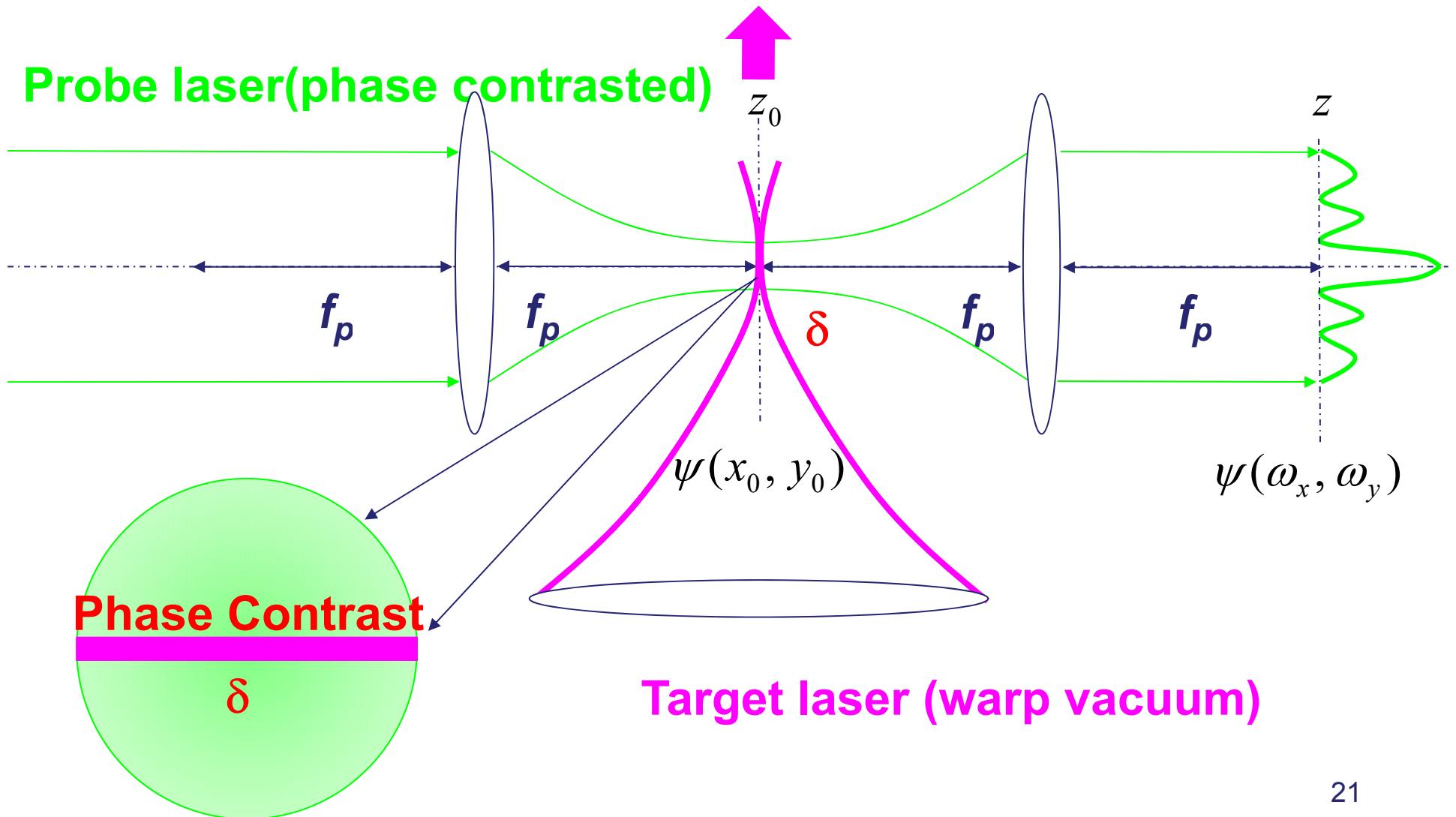


$$n_{\parallel} = 1 + \frac{16}{45} \frac{\alpha^2 U}{U_e}, \quad n_{\perp} = 1 + \frac{28}{45} \frac{\alpha^2 U}{U_e}$$

$$U_e = m_e^4 c^5 / \hbar^3 \approx 1.42 \times 10^6 \text{ J}/\mu\text{m}^3$$

ELI(~200J per ~20fs)
can reach $\Delta n \sim 10^{-9} \sim 10^{-10}$

Phase contrast imaging of vacuum



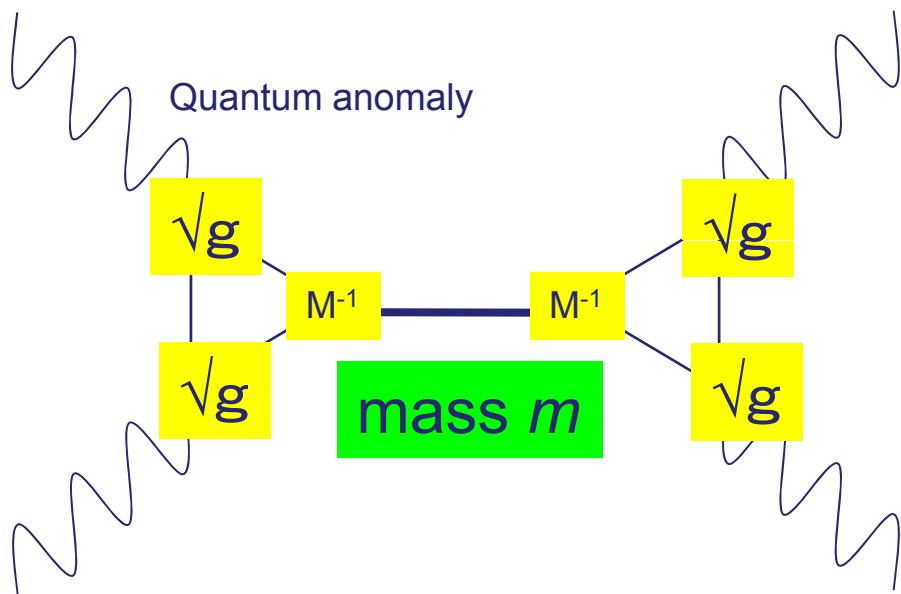
Beyond QED photon-photon interaction

$$L_{QED} = \frac{1}{360} \frac{\alpha^2}{m^4} [4(F_{\mu\nu}F^{\mu\nu})^2 + 7(F_{\mu\nu}\tilde{F}^{\mu\nu})^2]$$

$\phi F_{\mu\nu}F^{\mu\nu}$ $\sigma F_{\mu\nu}\tilde{F}^{\mu\nu}$

Away from $4 : 7 = \text{QCD}$, low-mass scalar ϕ , or pseudoscalar σ

Resonance in quasi-parallel collisions in low cms energy



If $M \sim M_{\text{Planck}}$, Dark Energy

$$gM^{-1}F^{\mu\nu}F_{\mu\nu}\phi$$

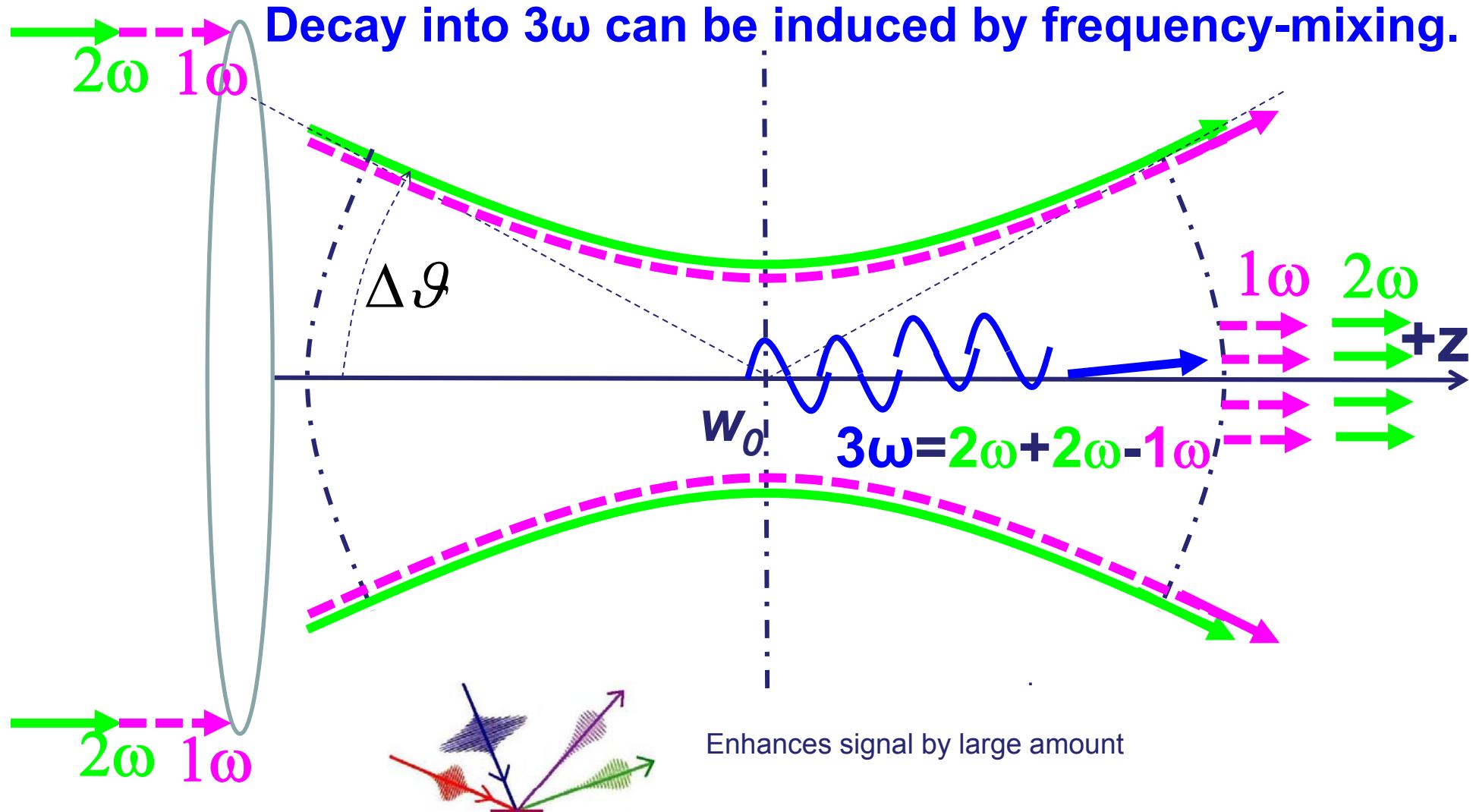
arXiv:1006.1762 [gr-qc]
Y. Fujii and K. Homma

QCD-instanton, Dark Matter

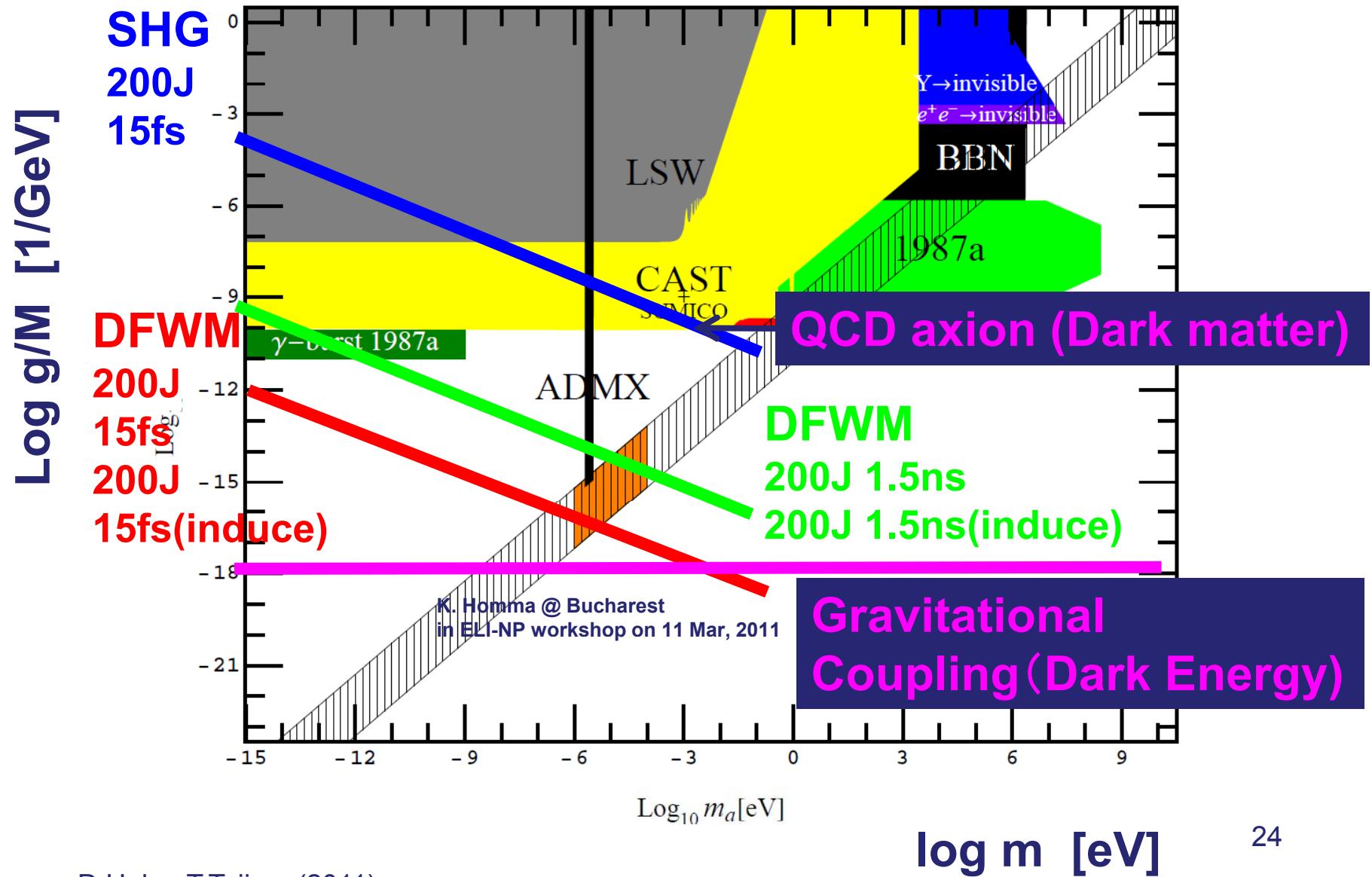
$$gM^{-1}F^{\mu\nu}\tilde{F}_{\mu\nu}\sigma$$

Degenerate Four-Wave Mixing (DFWM)

Laser-induced nonlinear optics in vacuum (cf. Nonlinear optics in crystal)



HFS road to unknown fields: dark matter and dark energy



Latest Development: CERN getting into the game

EuroNNac Workshop on novel accelerators (May 3-6, 2011)

EuCARD, EuroNNac Workshop, 3 - 6 May'11 / Programme

Tuesday 03 May 2011

Tuesday 03 May 2011

Introductory Presentations - Kjell Johnsen Auditorium (08:30-10:30)

- Conveners: Dr. Collier, Paul (CERN)

time	title	presenter
08:30	Goals of Network and Workshop (00h15')	ASSMANN, Ralph (CERN)
08:45	Accelerator R & D as Driver of Innovation (00h45')	HEUER, Rolf (CERN)
09:30	History and Outlook for Plasma Acceleration (00h30')	TOSHI, Tajima (LMU Munich)
10:00	Modern Lasers for Novel Acceleration Methods (00h30')	MOUROU, Gerard (ILE)

Coffee Break - 30-7-012 (10:30-11:00)

Introductory Presentations - Kjell Johnsen Auditorium (11:00-12:30)

- Conveners: Dr. Collier, Paul (CERN)

time	title	presenter
11:00	Accelerator R & D for Particle Physics (00h30')	MYERS, Steve (CERN)
11:30	Status Report Asia (00h30')	SHENG, Zhengming (Shanghai Jiao Tong University)
12:00	Status and Plans US (beam driven) (00h15')	HOGAN, Mark (SLAC)
12:15	Status and Plans US (Laser driven) (00h15')	ESAREY, Eric (LBNL)



IZEST

*International Center for
Zetta-Exawatt Science and Technology*

*Under the Aegis of
CEA, Ecole Polytechnique and
Ministry of Research and Education*

ELI (2010), now Mega Project on Extreme Laser (2011)

Extreme Light Infrastructure: EU decided (2010) at Czech, Hungary, and Romania
Now, Russia announced July 5, 2011: 6 Mega Projects (3-4B Euro) include **Extreme Laser**

Beyond Exawatt
Beyond 10kJ

ELI: serving Chair, Scientific Advisory Committee
Extreme Laser Mega Project (in budget negotiation):
Chief Scientific Advisor/
Mega Grant Honorary Director (suggested)
International team being formed:
IZEST (International Center for Zetawatt / Exawatt Science and Technology)

S&TRF
НАУКА И ТЕХНОЛОГИИ РФ

Евразийский открытый институт, используя обучение через интернет, реализует 18 программ...

По диаметру отверстия можно определить и вещества у...

05.07.11
Σ Стерлигов Иван
Правительственная комиссия по высоким технологиям и инновациям: Обсуждение
Обсуждение
Версия для печати
добавить ссылку

Сверхмощный лазер как интегратор науки

В числе [меганачальных проектов](#), которые будут реализованы на территории России, – Международный центр исследований экстремальных световых полей на основе сверхмощного лазерного комплекса в Нижнем Новгороде. Руководит центром всемирно известный физик [Жерар Мур](#) при поддержке Минобрнауки России. [STRF.ru](#) подробно рассказывал об этой работе в статье [«Российские учёные строят сверхмощный лазер»](#). Насколько значим этот проект для мировой науки, мы выяснили у [Тосики Тадзими](#), заведующего кафедрой физического факультета Университета Людвига Максимилиана в Мюнхене, председателя Международного комитета по сверхмощным лазерам ([International Committee on Ultra-High Intensity Lasers, ICUIL](#)).



Справка STRF.ru:
Международный комитет по сверхмощным лазерам – подразделение Международного союза фундаментальной и прикладной физики, основанное в 2003 году. Задача ICUIL – продвижение науки и технологии сверхмощных лазеров и координация исследований и разработок в этой области. Под сверхмощными лазерами в комитете понимают лазеры с интенсивностью 10^{19} ватт на см^2 и мощностью около 10 тераватт

На Ваш взгляд, что примечательного произошло в области сверхмощных лазеров в последнее время?

– Прошлый год стал эпохальным для нас благодаря решению Евросоюза о запуске проекта [Extreme Light Infrastructure](#) [ELI], включает целый ряд сверхмощных лазеров в нескольких регионах Европы], а также началу реальной работы [National Ignition Facility](#) в США – альтернативный токамакам проект термоядерной энергетики, основанный на лазерном нагреве и инерционном удержании плазмы. Мы предполагаем, что развитие сверхмощных лазеров и сопутствующих областей науки значительно ускорится, и стараемся способствовать

<http://strf.ru/>

XCELS (Russia)



Conclusions

- Optical approach: does it overtake the accelerator in high energy and fundamental physics?
- Collider physics requirements: == \rightarrow low density operation, **laser** with large energy per stage
- Energy frontier (beyond TeV) with precision w/ a few shots possible = non-collider paradigm of fundamental science
 - e.g. Lorentz invariance test , quantum gravity
- **High field science** approach: capability to explore new fields (dark matter; dark energy): SHG, DFWM, learning from NLO (in matter); zs metrology
- Join us at **IZEST**; collaboration btw ISTC and IZEST



The Cabin in Suzdal
(a Waka poem)

「遙か来ぬ ロシアの大地 夏残照
芳しきかな 丸太屋の宿」

“The distance I've come
far to the land of Russia
at summer's last cry
What fragrance and comfort
the logcabin lulls me in!”

Toshi Tajima
Sept, 2011

Спасибо! 30