

BSM problems w/ the  $r_p$  Problem

↳ PROTON RADIUS

Work "in progress" (on hold) w/ Gordon Krieger

→ looking for "COLLABORATORS"

↳ grad students to do calculations

REFERENCES

$r_p$  PROBLEM REVIEW; Pohl et al. 1301. 0905

New physics attempts: BTS & Yavin 1011. 4922

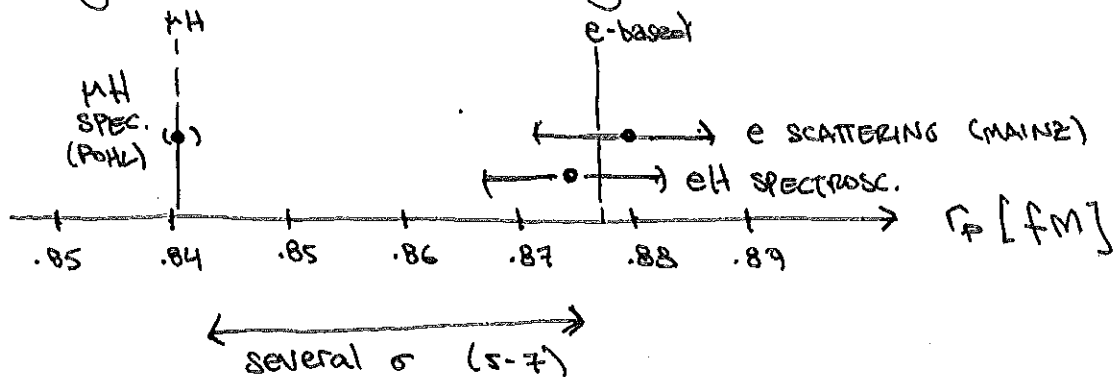
Barger et al. 1011. 3519

Carlson & Ridelw 1206. 3587

PI guys 1405. 4864

THIS TALK: ASSUME solution is new physics, see what we can do.

↳ (BIG ASSUMPTION)

Summary of the anomaly

$$r_p^e = 0.8770(45) \text{ fm}$$

$$r_p^\mu = 0.8409(04) \text{ fm}$$

ASIDE: WHAT IS  $r_p$ ? ← CHARGE RADIUS;  
DEVIATION FROM POINT-LIKE CHARGE

technically:  $r_p^2 \equiv -6 \frac{dG_E}{dQ^2} \Big|_{Q^2=0}$

← TRANSVERSE MOMENTUM

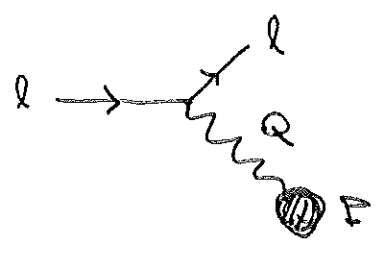
$G_E$ : SAATCHI'S ELECTRIC FORM FACTOR  
modifying POINT PROTON APPROX.

this is heuristically  $\langle r^2 \rangle = \int d^3r r^2 \psi(r)$   
but not quite → cf. Miller 1002.0357, 0908.1535  
IT'S REALLY A "TRANSVERSE CHARGE DENSITY"  
(ACCOUNTS FOR RECOIL EFFECTS)

↳ [ non-relativistically,  $\psi(\vec{r}_1, \vec{r}_2) = e^{i\vec{E} \cdot \vec{R}} \phi(\vec{r})$   
form factor is  $F = \int d^3r |\phi|^2 e^{i\vec{Q} \cdot \vec{r}/2}$

BUT RELATIVISTICALLY MUST SOLVE BS OR  $\psi$   
 $\psi$  DOESN'T FACTORIZE ↓

EFFECTIVELY:  $G_E(Q^2)$  IS PROB. AMP FOR  $p$  TO ABSORB  
THE EXCHANGED PHOTON (w/o EXCITING TO  
A NEW STATE)



↳ anyway, this is the actual thing that shows up in expts.

POINT PROTON  
↓

so:  $\delta V(\vec{r}) = -4\pi\alpha \int d^3q \frac{1}{q^2} [G_E(q^2) - 1] e^{-i\vec{q} \cdot \vec{r}}$

$\left( 1 - \frac{1}{6} q^2 r_p^2 + \mathcal{O}\left(\frac{r_p^4}{q^4}\right) \right)$

$= \frac{2\pi}{3} \alpha \delta^{(3)}(\vec{r})$       ↑

⇒  $\Delta E = \langle \psi_s | \delta V | \psi_s \rangle = \frac{2\pi}{3} \alpha |\psi_s(0)|^2 r_p^2$        $a_{\text{Bohr}} = \frac{1}{\alpha m_p}$

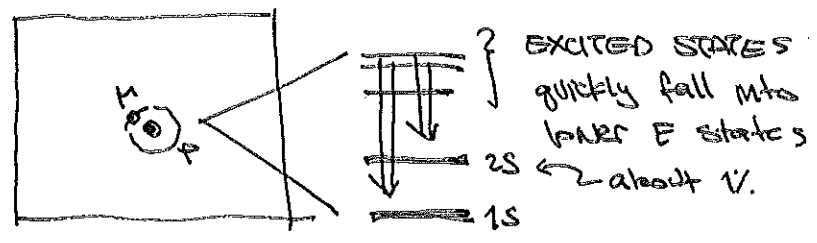
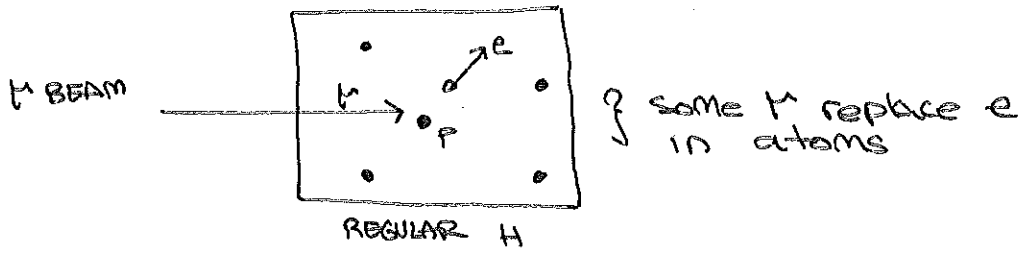
↑      ↑

need: -0.3 meV      shifts s-wave states  
≈  $(\frac{1}{a_{\text{Bohr}}})^3$

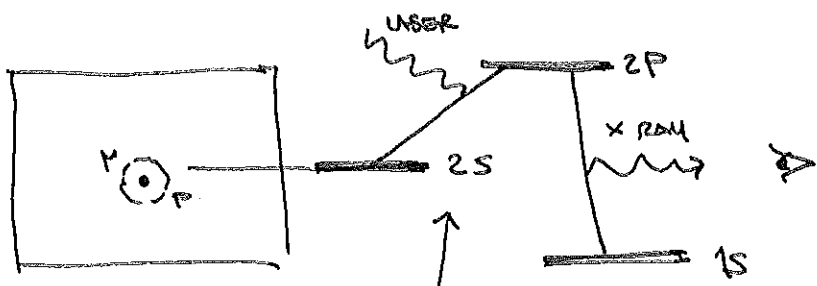
# What are we measuring

## I. SPECTROSCOPY → Atomic transitions

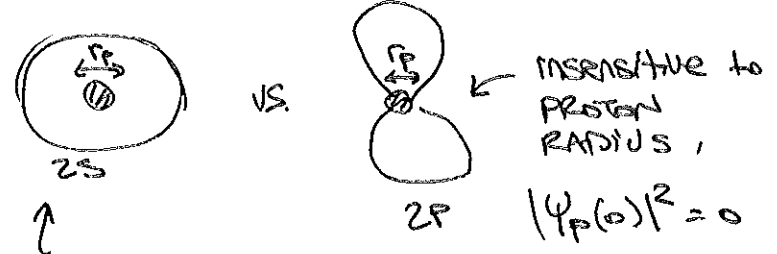
PREPARE MUONIC HYDROGEN



EXCITE 2S STATES INTO 2P STATES W/ LASER TUNED TO RIGHT FREQ.  
 Pohl had to find this!



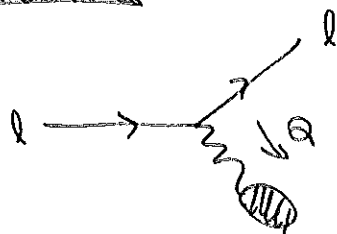
note: 2S state probes proton size while 2P does not.



So this ENERGY LEVEL SHIFTS DUE TO THE PROTON RADIUS

Similar meas. for e<sup>-</sup>

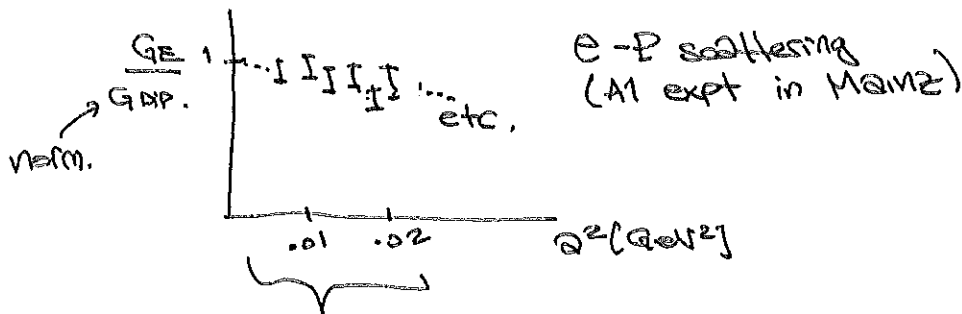
## II. SCATTERING



LOOKS LIKE DEEP INELASTIC SCATTERING ... BUT WE TAKE OPPOSITE LIMIT.

INSTEAD OF HIGH  $Q^2 \rightarrow$  SHORT WAVELENGTH PROBE P SUBSTRUCTURE  
 WANT LOW  $Q^2 \rightarrow$  LONG WAVELENGTH PROBE P "MACROSCOPE" PROPERTIES.

OF COURSE, CANNOT MEASURE  $Q^2 \rightarrow 0$   
 HAVE TO TAKE "AS LOW AS YOU CAN GO"  $Q^2 \rightarrow$  EXTRAPOLATE.



remark: in this ~linear region, linear extrapolation gives  $r_p \approx 0.84 \text{ fm}!!$

↳ BUT: in this regime,  $O(r_p^4)$  effects contribute if they are positive def. so 0.84 fm is UNDERSHOOTING.

[claim by Carl Carlson & Griffioen in  $r_p$  PROBLEM REVIEW]

Now ASSUME the resolution is BSM particle physics.

What ingredients do we need?

1. new particle to correct  $V_{CKM}$   $\rightarrow$  new mediator

$$\Delta V = (-)^{S+1} \left( \frac{g_e g_p}{e^2} \right) \alpha \frac{e^{-mr}}{r}$$

SPIN S  
mass r

2. COUPLES TO PROTONS

$\hookrightarrow$  couples to quarks/gluons  $\dashrightarrow$  also NEUTRONS?

3. COUPLES TO LEPTONS:  $\mu$  and/or  $e$

either pull  $\mu$  closer or push  $e$  further

IMMEDIATE CONCERN: LEPTON UNIVERSALITY

$\hookrightarrow$  flavor constraints in lepton sector

(so ignore SPIN-1)

but: this only applies to SPIN-1 mediators

basically: kinetic terms must be flavor diagonal s.t. rotation to fermion mass basis does not introduce tree-level FCNCs.

WE KNOW (but understated in sp literature) that SCALARS want to couple according to MASS.

$\hookrightarrow$  by assump. that scalar is SM singlet, it must couple to gauge singlet product of SM fields, eg  $\langle H \rangle L e$

in other words, couples  $\sim$  yukawas. so EXPECT MASS-WEIGHTED couplings.

$\Rightarrow$  this is exactly what we want!

$$\Delta E_{NP} = (-)^{S+1} \frac{\alpha}{2a^3} \frac{g_e g_p}{e^2} \frac{f(am)}{m^2} \leftarrow f(x) = \frac{x^4}{(1+x)^4}$$

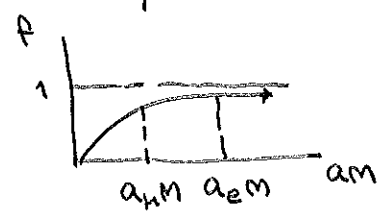
2S-2P

↑ see Jauch + Ray 1908.3536  
 ↗ BOHR RADIUS,  $a = \frac{1}{\alpha M_{red}} \approx \frac{1}{\alpha M_e}$

compare to proton radius correction to Lamb shift

$$\Delta E_{r_p} = \frac{2\alpha}{3n^3 a^3} \langle r_p^2 \rangle$$

$$\Rightarrow \Delta E_{TOTAL} = \frac{2\alpha}{3n^3 a^3} \left[ \underbrace{r_p^2}_{(\tilde{r}_p^2)_{EFFECTIVE}} + \left(\frac{3n^2}{2}\right) (-)^{S+1} \left(\frac{g_e g_p}{e^2}\right) \frac{f(am)}{m^2} \right]$$



SO: if  $(-)^{S+1} \left(\frac{g_e g_p}{e^2}\right)$  is universal & negative, eg for a spin-1 mediator,

then the effective  $r_p$  is SMALLER for electrons, vs. MUONS.

↑ this is the wrong direction!

WANT  $r_p^{(e)} > r_p^{(\mu)}$

⇒ disfavor spin-1 mediator

↑ can do "weird" things like  $U(1)_{E-H}$  BUT WE'LL NOT.

Alternative :  $|g_e| = |g_p|$  but opposite sign

eg  $(\mu-e)$  GAUGED

→ repulsive force for  $e-p$  system

BUT: THIS EFFECT IS SUPPRESSED @  
HIGH MOMENTUM TRANSFER, eg  
~~BE~~ FROM SCATTERING DATA.

→ would not explain Mainz  
proton radius measurements

SO: WE'RE LEFT w/ SPIN-0 COUPLING  $\sim$  YUKAWAS

$\hookrightarrow$  immediately think Higgs portal!

OPTIONS: SCALAR

✓

vs. PSEUDOSCALAR

↓

does not contribute in NR limit (derivative interaction) [cf 1011.3519]

SO EITHER REAL SPIN 0 or  $\sigma$  FIELD WORKS.

$$\text{LET } e^2 = \frac{g_{h\phi\phi}}{e^2} \Rightarrow \Delta\Gamma_P^2(\mu) = -6e^2 \frac{f(a_{\mu m})}{m^2}$$

$\uparrow$   
2/2

$$\Delta\Gamma_P^2(e) = -6e^2 \frac{m_e}{m_\mu} \frac{1}{m^2}$$

using  $f(a_{em}) \rightarrow 1$

$$\text{fit to } \Delta\Gamma_P^2(\mu) - \Delta\Gamma_P^2(e) = -0.063 \text{ fm}^2$$

$\uparrow$  SEE PI GUYS' PAPER  $\&$  notebook note: our  $e^2$  differs from theirs

BAULPARK ESTIMATE:

$$e^2 \sim 10^{-6}$$

CHECK THIS.

$\hookrightarrow$  nb PI GUYS: e-coupling

$$\frac{g_{e\phi\phi}}{e^2} \sim 1.3 \times 10^{-8}$$

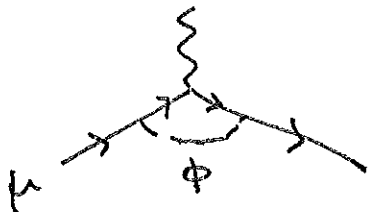
so now we have a target

$$\frac{g_{h\phi\phi}}{e^2} \sim 10^{-6}$$



IN FACT, IN "KICKSTARTER" FASHION, THERE'S ALSO AN IMMEDIATE "REACH" GOAL:

WE KNOW  $\mu$ -SPECIFIC (ie LEPTON NON-UNIVERSAL) FORCES CAN BE APPLIED TO THE  $(g-2)_\mu$  ANOMALY



- only depends on  $g_\mu$
- sensitive to both SPIN-0<sup>+</sup> AND  $[0^-]$

see Carlson & Ristau

Pseudoscalar: DIDN'T DO ANYTHING FOR  $\Gamma_P$  B/C DOESN'T AFFECT NON-RELATIVISTIC MEASUREMENT OF  $\Gamma_P$ .

BUT CONTRIBUTES @ SAME ORDER AS SCALAR IN  $(g-2)_\mu$

in fact, CONTRIBUTES W/ OPPOSITE SIGN  $\rightarrow$  DESTRUCTIVE INTERF.

We can actually use  $\Gamma_P$ 's.

IN FRACTIONAL TERMS,  $\Gamma_P$  DISCREPANCY IS  $\mathcal{O}(10^4)$  LARGER THAN  $(g-2)_\mu$  & ANY PARTICLE CONTRIBUTING TO THE LAMB SHIFT WILL ~~WE WILL YOU MIGHT~~ ALSO CONTRIBUTE TO  $(g-2)_\mu$ .

$\rightarrow$  SO YOU MIGHT NEED TO ENGINEER/TUNE A CANCELLATION.

~~analysis = we're getting ahead of ourselves  
WE HAVEN'T ESTABLISHED WHAT WE CAN SOLVE IF PROBLEM!~~

NOTE:

Carlson & Ristow :  $M \sim \mathcal{O}(100)$  MeV

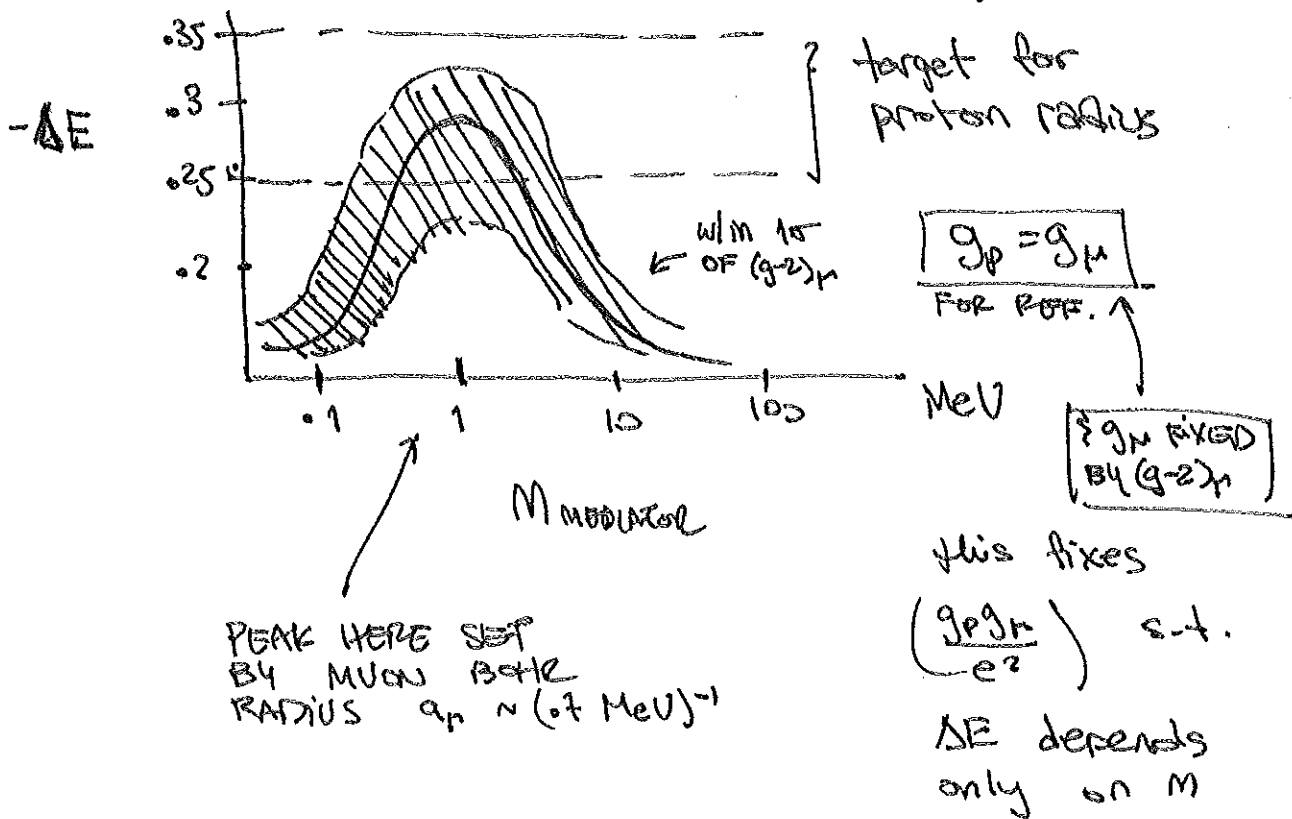
↳ require a tuning  
IN FACT, MAMBO reqs EVEN A FINE-TUNING.

DTS & Tray :  $M \sim \mathcal{O}(0.1) - \mathcal{O}(100)$  MeV

in this mass range  
 $M < M_\mu$  & target  
for muon coupling  $g_\mu$   
is insensitive to  $M$ .

for  $M \ll M_\mu \Rightarrow \boxed{\frac{g_\mu}{e} \sim 10^{-3}}$   
↑ SPIN-0

A NICE PLOT FOR ESTIMATES (DTS & TRAY)



PEAK HERE SET BY MUON BOHR RADIUS  $a_\mu \sim (0.7 \text{ MeV})^{-1}$

but we're getting ahead of ourselves

WE'VE ESTABLISHED WHERE WE WANT TO GO,  $(\frac{g_p g_n}{e^2}) \sim 10^{-6}$ , BUT NOW WE HAVE TO SEE WHAT STANDS IN OUR WAY.

1. BIGGEST CONSTRAINT: neutron-Pb scattering

↳ Barbieri & Ericson (1975) (Phys Lett 57B)

↑  
in fact, almost exact same story  
ANOMALY IN ATOMIC TRANSITIONS OF HIC ATOMS  
MAYBE NEW BOSON → YUKAWA POTENTIAL?

BUT: N SCATTERING ON HEAVY NUCLEI  
(in keV region for STUDYING N ELECTRIC  
POLARIZABILITY) GIVES A VERY  
SPECIFIC ANGULAR DISTRIBUTION  
COMING FROM THE STRONG FORCE

INTERFERENCE W/ A WEAKLY COUPLED  
SCALAR EXCHANGE DIAGRAM WILL  
MODIFY THE ANGULAR DISTRIBUTION.

Result:  $\frac{g_n^2}{4\pi} \left(\frac{\text{MeV}}{m}\right)^4 < 3 \times 10^{-11}$

↑  $\approx 1$  for us.

IN FACT: LIGHT HIGGS SEARCH  
FOR  $\phi \rightarrow e^+e^-$  PUSHES  
US TO  $M < 2M_e$

cf.  $\frac{g_p g_n}{e^2} \sim 10^{-6}$

SO CAN'T USE THIS  
FACTOR TO HELP?

so:  $\frac{e^2}{4\pi} \sim \frac{1}{137}$ ,  $4\pi \sim 10$ ,  $e^2 \sim 10$

TYPICALLY EXPECT  $g_p = g_n$

$10^{-7} \sim g_n g_p \sim g_n^2 < 10^{-9}$

↑ if  $g_n = g_p$ ... but this is preferred by  $(g_n)^2$   
•  $(\frac{g_n}{e}) (\frac{g_p}{e}) \sim 10^{-6}$  }  $(\frac{g_n}{e}) \sim 10^{-3}$

So this kinda sucks, by about  $O(100)$ .

WHAT TO DO: make  $g_p = g_n$  SMALL

↓

make  $g_n$  LARGE

BUT THEN  $(g-2)\mu$  BECOMES  
A BOUND, NOT AN OPPORTUNITY

(ie you overshoot)

BUT I'M FROM UCI, SO MAYBE I CONSIDER ISASPIN,  
 $g_p \neq g_n$  in such a way that  $g_p \gg g_n$ .

↑ we'll explore this.

### OTHER BOUNDS

2.  $(g-2)\mu$  ← already mentioned  $(g-2)\mu$

$(g-2)e$  IS ACTUALLY USED TO DEFINE  $\alpha$

↳ SHIFT IN  $(g-2)e$  → SHIFT IN  $\alpha$

$$\Delta\alpha = 2\pi \Delta Q_e$$

↑ CHECK VS. Rb & Cs  
ATOM CHECKS.  
SEE REFS IN DTS & ICM.

↑ THIS IS  $(g-2)e/2$   
NOT Bohr RADIUS

$$\text{RESULT: } \frac{\alpha}{2\pi} \left(\frac{g_e}{e}\right)^2 \int_0^1 \frac{(1-z)^2(1+z)}{(1-z)^2 + (m/m_e)z} dz < \frac{\alpha}{2\pi} 15 \times 10^{-9}$$

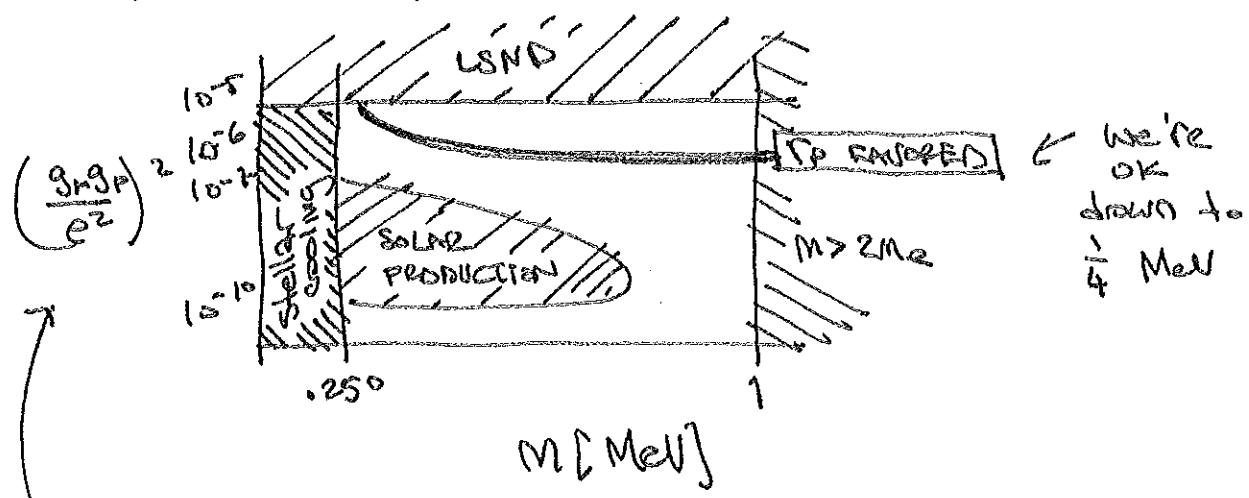
↳

→  $3/2$

$$\Rightarrow \left[ \frac{g_e}{e} \lesssim 10^{-4} \right] \text{ for } M \sim \text{MeV}$$

→ also  $g_n^2 \lesssim 10^6$  from  $\left(\frac{M_p}{\Delta e}\right)$  factors

### 3. ASTROPHYSICS & LIGHT PARTICLE BOUNDS from PI GUYS



nb I RECALCULATED THEIR BOUNDS, WHICH WERE CAST IN TERMS OF  $(g/g_p)$ .

### 4. others: PLAXIC (KARM decays to $\mu$ ) OTHER ATOMIC SYSTEMS w/ $\mu$ .

↑  
anyway, 1-3 are the big ones.

Now get back to model building

So I WANT: ① HIGGS PORTAL → nonuniversal couplings

w/  $m(\frac{1}{2}, 1)_{\text{Higgs}}$

$$\frac{g_p g_M}{e^2} \sim 10^{-6}$$

$$\frac{g_n}{4\pi} < 10^{-11}$$

↑  
SOMETHING LIKE  
 $\phi^2 H^2$  & RELATED  
TERMS

↘ also automatically  
gives couplings to  
quarks & leptons

② ISOSPIN VIOLATION →  $g_n \ll g_p$   
TO AVOID  $n$ -PB SCATTERING  
BOUNDS.

simplest Higgs portal doesn't work

↳ can only really play w/ mixing w/ Higgs

↳ cannot further discriminate  
between  $g_p, g_n$  &  $g_n$   
ISOSPIN INDEP.

RECALL COUPLING OF SCALAR (HIGGS) TO NUCLEONS  
(SHIFMAN et al, HIGGS LOW-E QM — SEE  
DIRECT DETECTION REVIEWS)

$$\sum_{q \text{ quarks}} C_q M_q \phi \bar{q} q \rightarrow C_N M_N \phi \bar{N} N \quad (\text{assume } C_q \approx 0)$$

$$\uparrow$$

$$C_N \equiv \sum_{q \text{ quarks}} C_q f_q^{(N)} + \frac{2}{2B} f_B^{(N)} \quad (\sum_{q \text{ quarks}} C_q)$$

only this has some  
discrimination between  
proton & neutron.

NUCLEAR FORM FACTORS

DARKSUSY } eg  $f_u^{(p)} = 0.023$      $f_u^{(n)} = 0.019$   
 $f_d^{(p)} = 0.034$      $f_d^{(n)} = 0.041$

CAN SEE THAT W'LL NEED TO TUNE  $C_q$ 'S A BIT.

other options

PROBLEM IS THAT IN  $C_g M_g \phi \tilde{g} \tilde{g} + C_e M_e \phi \tilde{e} \tilde{e}$

$C_g = C_e = \text{UNIVERSAL}$

so to discriminate P from N, need to be able to tune  $c_u$  vs  $c_d$ .

I. More hidden sector ~~(UNIVERSAL)~~

~~(Higgs Portal)~~ ~~POSSIBLE~~ HIDDEN HIGGS

INTRODUCE INERT (no vev) HIGGS  $\tilde{H}$  (DOUBLET) WHICH ONLY COUPLES TO, SM, UP-TYPES.

~~CAN YOU STILL GET THE~~  
THEN MIX  $\phi$  WITH  $\tilde{H}$  TO GET A DIFFERENT  $\phi$  COUPLING TO UP QUARKS

BUT: Flavor constraints

YOU HAVE TO ENGINEER (OR ASSUME) THAT  $\tilde{H}$  HAPPENS TO ALSO COUPLE PROPORTIONALLY TO THE SM YUKAWAS.

II MORE VISIBLE SECTOR

↳ MORE HIGGS. eg ZADM IS WELL MOTIVATED.

→ SPENT LOTS OF TIME PLAYING W/ THIS IN ASEN, ENDS UP NOT WORKING, WE THINK.



because we weren't clever enough to diagonalize the 3x3 subject to our constraints.

{DEF. DOESN'T WORK IN THE SMALL MIXING LIMIT WHICH IS THE OBVIOUS FIRST ATTEMPT}

## SOME ISSUES W/ ZHDM

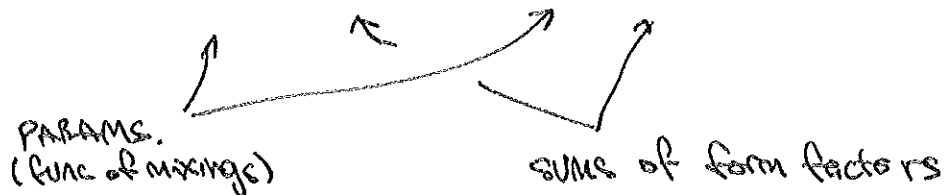
→ PHENO: HEAVY HIGGS SEARCHES CONSTRANED  
PARAMETER SPACE

→ ~~TARGET~~ CONSTRAINTS

IN THE SMALL MIXING LIMIT, END UP W/

$$C_p = d_u (0.14) + d_d (0.23)$$

$$C_n = d_u (0.14) + d_d (0.24)$$



HAVE TO TUNE AT LEAST TO 1% TO CANCEL  $C_n$

then you also know that  $C_R = d_d \frac{M_R}{M_n}$   
↑  
↑ for type 4

these have bounds from  
 $(9-2)e, M$

SO YOU CAN'T JUST TUNE THE MIXINGS IN  
THE SQUARES TO CANCEL  $C_n$ , THEN  
PUMP UP  $C_p$  TO HIT TP TARGET,

B/C THIS "PUMPING UP" ALSO PUSHES  $C_p$  TOO LARGE!

I CAN TRY OTHER ZHDM, BUT IF U'D TALK TO  
ONE HIGGS, NO ROOM TO TUNE AWAY  $C_n$ .

IF I  $\neq$  U, SAME PROB. AS ABOVE.



## So what now?

PROJECT SHOWED ETC OTHER IDEAS CAME UP.  
BUT WE HAVE A LOT OF WORK ALREADY DONE  
(MATHEMATICS NOTEBOOKS, DRAFT)

→ WE'RE LOOKING FOR COLLABORATORS (GRAD STUDENTS?)  
INTERESTED IN PLAYING WITH THIS.

## Ways forward

1. LARGE MIXING IN 2HDM  
↳ MIXING COULD GIVE CANCELLATIONS  
(THIS IS WHAT HAPPENS IN 1HDM  
→ IN SEE-SAW)
  2. 3HDM? still 3 nontrivial benchmarks ...
  3. Bigger hidden sector?
  4. connection to Dark Matter?  
↳ WE'VE BASICALLY BEEN RIPPING ON  
THE HIGGS PORTAL.  
CAN WE USE  $\phi$  AS A MEDIATOR?
- ! others?

## ONGOING EXPTS

- MUSE: (2016 start??)  $p$ - $p$  SCATTERING
- other  $h^0$  ATOMS (deuterons?)